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# Supersymmetric codimension-two branes and $U(1)_R$ mediation in 6D gauged supergravity

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ABSTRACT: We construct a consistent supersymmetric action for brane chiral and vector multiplets in a six-dimensional chiral gauged supergravity. A nonzero brane tension can be accommodated by allowing for a brane-localized Fayet-Iliopoulos term proportional to the brane tension. When the brane chiral multiplet is charged under the bulk  $U(1)_R$ , we obtain a nontrivial coupling to the extra component of the  $U(1)_R$  gauge field strength and a singular scalar self-interaction term. Dimensionally reducing to 4D on a football supersymmetric solution, we discuss the implication of such interactions for obtaining the  $U(1)_R$  D-term in the 4D effective supergravity. By assuming the bulk gaugino condensates and nonzero brane F- and/or D-term for the uplifting potential, we have all the moduli stabilized with a vanishing cosmological constant. The brane scalar with nonzero R charge then gets a soft mass of order the gravitino mass. The overall sign of the soft mass squared depends on the sign of the R charge as well as whether the brane F- or D-term dominates.

KEYWORDS: Field Theories in Higher Dimensions, Supersymmetry Breaking, Supergravity Models, Supersymmetric Standard Model.

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# 1. Introduction

Supersymmetry(SUSY) [1] has been around us as one of the most promising candidates for physics beyond the Standard Model. When the SUSY breaking occurs at the electroweak scale, it can be a solution to the gauge hierarchy problem, thanks to the cancellation of the quadratic divergences to the Higgs mass between the SM particles and their superpartners. If the SUSY breaking is parametrized in terms of the soft mass parameters that respect the SM gauge symmetry, over 100 additional parameters would lead to unacceptably large FCNCs and CP violations. In gravity mediation [2] where the hidden sector SUSY breaking is transmitted to the visible sector by gravity only, the weak scale soft mass parameters can be naturally generated. However, the almost flavor-universal soft masses are unexplained in this context because one cannot forbid the flavor-dependent contact interactions between the visible and hidden sectors by any known symmetry. Therefore, several alternative mechanisms of SUSY mediation giving the flavor-universal soft masses have been suggested and discovered: gauge mediation [3], anomaly mediation [4], gaugino mediation [5], etc.

For recent years, there have been a plenty of interest in the flux compactifications for the SUSY phenomenology, in particular, in the context of the KKLT-type compactifications [6], due to the fact that fluxes combined with non-perturbative effect can fix all the moduli of the extra dimensions endowed from string theory and the soft mass parameters have a distinct pattern compared to the ones obtained in the other SUSY mediations, that leads to the so called anomaly-modulus mixed mediation or the mirage mediation [7]. In the present era that the Large Hadron Collider(LHC) is turning on soon, it is compelling and very important to identify the distinguishable features of the reasonable and accessible SUSY mediation mechanisms.

The model that we are considering is the Salam-Sezgin supergravity [8] where the  $U(1)_R$  subgroup of the bulk R symmetry is gauged by a bulk vector multiplet. Due to the  $U(1)_R$  gauging, there appears a nonzero positive scalar potential for the dilaton, in contrast to the 5D gauged supergravity. In this model, due to the cancellation between the dilaton potential and the  $U(1)_R$  gauge flux, the 4D Minkowski space with factorizable extra dimensions compactified on a sphere was obtained. The solution has been generalized rather recently to the unwarped or warped 4D Minkowski solutions with nonzero brane tensions [9-12], with the hope of achieving a self-tuning of the cosmological constant on a codimension-two brane [13]. In this case, nonzero brane tensions need to be introduced at the conical singularities that are caused by the deficit angles. The brane tensions are, however, regarded as breaking the bulk SUSY explicitly at the action level. Therefore, most recently, the SUSY action for the case with a brane tension has been constructed by adding localized FI terms and localized corrections to the Chern-Simons term in the field strength for the Kalb-Ramond field as well as by modifying the fermionic SUSY transformations [14]. Consequently, it has been shown that the unwarped football solution with arbitrary brane tensions is a new SUSY background solution preserving the 4D  $\mathcal{N} = 1$  SUSY [14].

In this paper, by extending the previous result for the SUSY action for a brane tension action of ref. [14], we first construct a consistent SUSY action in the presence of brane multiplets in the Salam-Sezgin supergravity by following the Noether method undertaken in 6D ungauged supergravity [15]. In this process, we need to add the localized terms depending on the brane multiplets to the field strengths and modify the fermionic SUSY transformations and the SUSY/gauge transformations of the Kalb-Ramond field. Consequently, as in the ungauged supergravity [15], we find that the kinetic term for the brane chiral multiplet has a nontrivial dilaton coupling while the gauge kinetic function for the brane vector multiplet is trivial. Moreover, the brane-induced F- and D-terms have nontrivial moduli couplings. When the brane chiral multiplet is charged under the  $U(1)_R$ , we also obtain a nontrivial coupling of the brane scalar to the gauge field strength as well as a singular scalar self interaction.

For the SUSY background solution with football-shaped extra dimensions, we consider the low energy action with light bulk and brane modes in the 4D effective supergravity. The  $U(1)_R$  gauge symmetry appears anomalous as the bulk Green-Schwarz counterterm generates the 4D  $U(1)_R$  anomalies for a nonzero gauge flux and a localized FI term. Moreover, the bulk  $U(1)_R$  vector multiplet gets a mass of order the 4D Planck scale by a Green-Schwarz mechanism. From the effective  $U(1)_R$  D-term potential, we have fixed only one of two moduli, i.e. the T modulus, due to the interplay between the field-dependent and constant  $U(1)_R$  FI terms. By assuming that there appears an additional scalar potential due to two bulk gaugino condensates generated below the compactification scale in an extend bulk theory and consequently adding the brane F- and/or D-term as the uplifting potential, we show that the remaining S modulus is also stabilized at the Minkowski vacuum. Due to the shift of the VEV of the T modulus from the one obtained only by the  $U(1)_R$  D-term, we show that a brane scalar field with nonzero R charge can get a nonzero soft mass in comparable size to the gravitino mass. We dub this new possibility " $U(1)_R$  mediation". The brane scalar soft mass does not depend on the  $U(1)_R$  gauge coupling because the mass of the  $U(1)_R$  vector multiplet is also proportional to the  $U(1)_R$  gauge coupling. The overall sign of the soft mass depends on whether the brane F-term or D-term dominates.

The paper is organized as follows. We start with describing the bulk Salam-Sezgin supergravity and then present the SUSY action for the chiral and vector multiplets living on the conical branes and the brane-localized D- and F-terms with nontrivial moduli dependence. Next we review on the recently found SUSY brane solution and continue to derive the low energy effective action for light bulk and brane fields in the same SUSY brane background to identify the corresponding 4D effective supergravity. We also discuss on the moduli stabilization and the SUSY breaking in the presence of the bulk gaugino condensates and nonzero brane F- and/or D-term uplifting potentials. Finally, the conclusion is drawn.

## 2. Model setup

The six-dimensional Salam-Sezgin supergravity [8] consists of gravity coupled to a dilaton field  $\phi$ , a Kalb-Ramond(KR) field  $B_{MN}$ , along with the SUSY fermionic partners, the gravitino  $\psi_M$ , the dilatino  $\chi$ . Moreover, it also contains a bulk  $U(1)_R$  vector multiplet  $(A_M, \lambda)$  that gauges the *R*-symmetry of six-dimensional supergravity. All the bulk fermions are 6D Weyl.

The complete bulk Langrangian [8] is given (up to four fermion terms) by

$$e_{6}^{-1}\mathcal{L}_{\text{bulk}} = R - \frac{1}{4}(\partial_{M}\phi)^{2} - \frac{1}{12}e^{\phi}G_{MNP}G^{MNP} - \frac{1}{4}e^{\frac{1}{2}\phi}F_{MN}F^{MN} - 8g^{2}e^{-\frac{1}{2}\phi} \\ + \bar{\psi}_{M}\Gamma^{MNP}\mathcal{D}_{N}\psi_{P} + \bar{\chi}\Gamma^{M}\mathcal{D}_{M}\chi + \bar{\lambda}\Gamma^{M}\mathcal{D}_{M}\lambda \\ + \frac{1}{4}(\partial_{M}\phi)(\bar{\psi}_{N}\Gamma^{M}\Gamma^{N}\chi + \bar{\chi}\Gamma^{N}\Gamma^{M}\psi_{N}) \\ + \frac{1}{24}e^{\frac{1}{2}\phi}G_{MNP}(\bar{\psi}^{R}\Gamma_{[R}\Gamma^{MNP}\Gamma_{S]}\psi^{S} + \bar{\psi}_{R}\Gamma^{MNP}\Gamma^{R}\chi \\ - \bar{\chi}\Gamma^{R}\Gamma^{MNP}\psi_{R} - \bar{\chi}\Gamma^{MNP}\chi + \bar{\lambda}\Gamma^{MNP}\lambda) \\ - \frac{1}{4\sqrt{2}}e^{\frac{1}{4}\phi}F_{MN}(\bar{\psi}_{Q}\Gamma^{MN}\Gamma^{Q}\lambda + \bar{\lambda}\Gamma^{Q}\Gamma^{MN}\psi_{Q} + \bar{\chi}\Gamma^{MN}\lambda - \bar{\lambda}\Gamma^{MN}\chi) \\ + i\sqrt{2}ge^{-\frac{1}{4}\phi}(\bar{\psi}_{M}\Gamma^{M}\lambda + \bar{\lambda}\Gamma^{M}\psi_{M} - \bar{\chi}\lambda + \bar{\lambda}\chi).$$
(2.1)

The field strengths of the gauge and Kalb-Ramond(KR) fields are defined as

$$F_{MN} = \partial_M A_N - \partial_N A_N, \qquad (2.2)$$

$$G_{MNP} = 3\partial_{[M}B_{NP]} + \frac{3}{2}F_{[MN}A_{P]}, \qquad (2.3)$$

and satisfy the Bianchi identities

$$\partial_{[Q}F_{MN]} = 0, \tag{2.4}$$

$$\partial_{[Q}G_{MNP]} = \frac{3}{4}F_{[MN}F_{QP]}.$$
(2.5)

For  $\delta_{\Lambda}A_M = \partial_M \Lambda$  under the U(1)<sub>R</sub>, the field strength for the KR field is made gauge invariant by allowing for  $B_{MN}$  to transform as

$$\delta_{\Lambda} B_{MN} = -\frac{1}{2} \Lambda F_{MN}. \tag{2.6}$$

All the spinors have the same R charge +1, so the covariant derivative of the gravitino, for instance, is given by

$$\mathcal{D}_M \psi_N = (\partial_M + \frac{1}{4} \omega_{MAB} \Gamma^{AB} - igA_M) \psi_N.$$
(2.7)

The local  $\mathcal{N} = 2$  SUSY transformations are (up to trilinear fermion terms):

$$\delta e_M^A = -\frac{1}{4} \bar{\varepsilon} \Gamma^A \psi_M + \text{h.c.}, \qquad (2.8)$$

$$\delta\phi = \frac{1}{2}\bar{\varepsilon}\chi + \text{h.c.},\tag{2.9}$$

$$\delta B_{MN} = A_{[M} \delta A_{N]} + \frac{1}{4} e^{-\frac{1}{2}\phi} (\bar{\varepsilon} \Gamma_M \psi_N - \bar{\varepsilon} \Gamma_N \psi_M + \bar{\varepsilon} \Gamma_{MN} \chi + \text{h.c.}), \qquad (2.10)$$

$$\delta\chi = -\frac{1}{4}(\partial_M\phi)\Gamma^M\varepsilon + \frac{1}{24}e^{\frac{1}{2}\phi}G_{MNP}\Gamma^{MNP}\varepsilon, \qquad (2.11)$$

$$\delta\psi_M = \mathcal{D}_M \varepsilon + \frac{1}{48} e^{\frac{1}{2}\phi} G_{PQR} \Gamma^{PQR} \Gamma_M \varepsilon, \qquad (2.12)$$

$$\delta A_M = \frac{1}{2\sqrt{2}} e^{-\frac{1}{4}\phi} (\bar{\varepsilon}\Gamma_M \lambda + \text{h.c.}), \qquad (2.13)$$

$$\delta\lambda = \frac{1}{4\sqrt{2}} e^{\frac{1}{4}\phi} F_{MN} \Gamma^{MN} \varepsilon - i\sqrt{2}g \, e^{-\frac{1}{4}\phi} \varepsilon.$$
(2.14)

The above spinors are chiral with handednesses

$$\Gamma^7 \psi_M = +\psi_M, \quad \Gamma^7 \chi = -\chi, \quad \Gamma^7 \lambda = +\lambda, \quad \Gamma^7 \varepsilon = +\varepsilon.$$
 (2.15)

Taking into account that  $\Gamma^7 = \sigma^3 \otimes \mathbf{1}$  (see appendix A), the 6D (8-component) spinors can be decomposed to 6D Weyl (4-component) spinors as

$$\psi_M = (\tilde{\psi}_M, 0)^T, \quad \chi = (0, \tilde{\chi})^T, \quad \lambda = (\tilde{\lambda}, 0)^T, \quad \varepsilon = (\tilde{\varepsilon}, 0)^T.$$
 (2.16)

For later use, we decompose the 6D Weyl spinor  $\tilde{\psi}$  to  $\tilde{\psi} = (\tilde{\psi}_L, \tilde{\psi}_R)^T$ , satisfying  $\gamma^5 (\tilde{\psi}_L, 0)^T = +(\tilde{\psi}_L, 0)^T$  and  $\gamma^5 (0, \tilde{\psi}_R)^T = -(0, \tilde{\psi}_R)^T$ . Henceforth we drop the tildes for simplicity.

We can show that the action for the Lagrangian (2.1) is invariant under the above SUSY transformations up to the trilinear fermion terms and the Bianchi identities as follows,

$$\delta \mathcal{L}_{\text{bulk}} = e_6 \bigg[ -\frac{1}{24} e^{\frac{1}{2}\phi} \bigg( \partial_S G_{MNP} - \frac{3}{4} F_{MN} F_{SP} \bigg) \bigg( \bar{\psi}_R \Gamma^{RMNPS} \varepsilon - \bar{\chi} \Gamma^{SMNP} \varepsilon + \text{h.c.} \bigg) + \frac{1}{4\sqrt{2}} e^{\frac{1}{4}\phi} \bigg( \partial_Q F_{MN} \bar{\lambda} \Gamma^{QMN} \varepsilon + \text{h.c.} \bigg) \bigg].$$
(2.17)

Thus, as will be seen later, the SUSY variation of the brane action can be cancelled with the bulk variation (2.17) by modifying the Bianchi identities (2.4) and (2.5).

## 3. Supersymmetric codimension-two brane actions

We introduce a chiral multiplet with nonzero *R*-charge and a vector multiplet on the brane. Then, by adopting the Noether method, we construct a consistent SUSY action for the brane multiplets that is invariant under the modified bulk SUSY variations. As a result, we show that the brane multiplets have nontrivial couplings to the bulk fields through the modified field strengths. We also consider a supersymmetric brane-localized gravitino mass term.

## **3.1** The $Z_2$ orbifold parities

In order to project out half the bulk supersymmetries on the brane and define an  $\mathcal{N} = 1$  brane SUSY, we assume an orbifold  $Z_2$  symmetry around each brane. If the local complex coordinate around the brane is z (in locally polar coordinates  $z = re^{i\theta}$ ), then the  $Z_2$  symmetry corresponds to

$$z \leftrightarrow -z$$
 (or  $\theta \leftrightarrow \theta + \pi$ ). (3.1)

In the case with two branes system, the warped vacua of [11] have an axially symmetric internal space. The above  $Z_2$  symmetry about both branes present, is just a discrete subgroup of the axial symmetry. On the other hand, for the general warped solutions with multiple branes [12], we require the holomorphic function V(z) in the metric to satisfy the condition  $|V(-z + z_i)| = |V(z - z_i)|$ , where  $z_i$  is the *i*-th brane position.

We should then assign  $Z_2$  parities to all bulk fields as well as the SUSY variation parameters  $\varepsilon_L$  and  $\varepsilon_R$ . Being consistent with the bulk action and the SUSY transformation, we make a choice of parities for the fields and the SUSY variation parameter as

even : 
$$\psi_{\alpha L}$$
,  $\psi_{aR}$ ,  $\lambda_L$ ,  $\chi_R$ ,  $\varepsilon_L$ ,  $A_{\alpha}$ ,  $B_{\alpha\beta}$ ,  $B_{ab}$ ,  $\phi$ , (3.2)

odd : 
$$\psi_{\alpha R}, \psi_{aL}, \lambda_R, \chi_L, \varepsilon_R, A_a, B_{\alpha a}.$$
 (3.3)

where the gauge field, the Kalb-Ramond field and the gravitino have been written with locally flat indices, e.g.,  $A_A = e_A^M A_M$ , so that the parity assignments do not depend on the coordinate system. It is obvious that the above choice of parities forces  $\varepsilon_R$  to vanish on the brane position. Henceforth we denote the 4D Weyl spinor of each bulk fermion surviving on the brane by  $\lambda_+$  and  $\varepsilon_+$ , etc, satisfying  $\gamma^5 \lambda_+ = \lambda_+$  and  $\gamma^5 \varepsilon_+ = \varepsilon_+$ , etc.

## 3.2 The supersymmetric action for brane multiplets

We consider a nonzero brane tension as well as brane matter multiplets: a brane chiral multiplet<sup>1</sup>  $(Q, \psi_Q)$ , the superfield of which has an R charge -r, and a brane vector multiplet  $(W_{\mu}, \Lambda)$ . Then, by employing the Noether method for the local SUSY, we find that the supersymmetric action for the bulk-brane system (up to four fermion terms) is composed of the original bulk action (2.1) with the field strength tensors  $G_{MNP}$  and  $F_{MN}$  being replaced by the modified ones  $\hat{G}_{MNP}$  and  $\hat{F}_{MN}$ , respectively, and the brane action as follows,

$$\mathcal{L} = \mathcal{L}_{\text{bulk}}(G \to \hat{G}, F \to \hat{F}) + \delta^2(y)\mathcal{L}_{\text{brane}}$$
(3.4)

with

$$\mathcal{L}_{\text{brane}} = e_4 \bigg[ e^{\frac{1}{2}\phi} \bigg( - (D^{\mu}Q)^{\dagger} D_{\mu}Q + \frac{1}{2} \bar{\psi}_Q \gamma^{\mu} D_{\mu} \psi_Q + \text{h.c.} \bigg) \\ + \sqrt{2} irg e^{\frac{1}{4}\phi} \bar{\psi}_Q \lambda_+ Q + \text{h.c.} - 4rg^2 |Q|^2 - T \\ + e^{\frac{1}{2}\phi} \bigg( \frac{1}{2} \bar{\psi}_{\mu+} \gamma^{\nu} \gamma^{\mu} \psi_Q (D_{\nu}Q)^{\dagger} + \frac{1}{2} \bar{\psi}_Q \gamma^{\mu} \chi_+ D_{\mu}Q + \text{h.c.} \bigg) \\ - \frac{1}{4} W_{\mu\nu} W^{\mu\nu} + \frac{1}{2} \bar{\Lambda} \gamma^{\mu} D_{\mu} \Lambda + \text{h.c.} \\ - ie \sqrt{2} e^{\frac{1}{2}\phi} Q \bar{\psi}_Q \Lambda + \text{h.c.} - \frac{1}{2} e^2 |Q|^4 e^{\phi} \\ - \frac{1}{4\sqrt{2}} \bar{\Lambda} \gamma^{\mu} \gamma^{\nu\rho} \psi_{\mu+} W_{\nu\rho} - \frac{i}{2\sqrt{2}} e |Q|^2 e^{\frac{1}{2}\phi} \bar{\Lambda} \gamma^{\mu} \psi_{\mu+} + \text{h.c.} \\ - \frac{i}{\sqrt{2}} e |Q|^2 e^{\frac{1}{2}\phi} \bar{\chi}_+ \Lambda + \text{h.c.} \bigg].$$
(3.5)

The SUSY transformations of the brane chiral multiplet are

$$\delta Q = \frac{1}{2}\bar{\varepsilon}_{+}\psi_{Q}, \quad \delta\psi_{Q} = -\frac{1}{2}\gamma^{\mu}\varepsilon_{+}D_{\mu}Q. \tag{3.6}$$

On the other hand, the SUSY transformations of the brane vector multiplet are

$$\delta W_{\mu} = \frac{1}{2\sqrt{2}} \bar{\varepsilon}_{+} \gamma_{\mu} \Lambda + \text{h.c.}, \qquad (3.7)$$

$$\delta\Lambda = \frac{1}{4\sqrt{2}}\gamma^{\mu\nu}\varepsilon_+ W_{\mu\nu} + \frac{i}{2\sqrt{2}}e|Q|^2 e^{\frac{1}{2}\phi}\varepsilon_+.$$
(3.8)

Here the brane gauge field strength is  $W_{\mu\nu} = \partial_{\mu}W_{\nu} - \partial_{\nu}W_{\mu}$  and the covariant derivatives of the brane multiplets are

$$D_{\mu}Q = (\partial_{\mu} + irgA_{\mu} - ieW_{\mu})Q, \qquad (3.9)$$

$$D_{\mu}\psi_{Q} = (\partial_{\mu} + i(r-1)gA_{\mu} - ieW_{\mu} + \frac{1}{4}\omega_{\mu\alpha\beta}\gamma^{\alpha\beta})\psi_{Q}, \qquad (3.10)$$

$$D_{\mu}\Lambda = (\partial_{\mu} - igA_{\mu} + \frac{1}{4}\omega_{\mu\alpha\beta}\gamma^{\alpha\beta})\Lambda.$$
(3.11)

<sup>&</sup>lt;sup>1</sup>The 4D chirality of the fermion in the brane chiral multiplet is taken to be right-handed in contrast to the  $Z_2$ -even gravitino and the  $Z_2$ -even gaugino and the brane gaugino. So, the conventional chiral superfield containing a left-handed fermion,  $(Q^*, (\psi_Q)^c)$ , should have an opposite R charge, namely, r for  $Q^*$  and r-1 for  $(\psi_Q)^c$ .

We note that the R charges of the component fields in the brane chiral multiplet are different by +1 as known to be the case in 4D local SUSY [18]. The gaugino of a brane vector multiplet also has the same R charge +1 as the bulk gravitino.

The modified field strength tensors are

$$\hat{G}_{\mu m n} = G_{\mu m n} + \left(J_{\mu} - \xi A_{\mu}\right) \epsilon_{m n} \frac{\delta^2(y)}{e_2},$$
(3.12)

$$\hat{G}_{\tau\rho\sigma} = G_{\tau\rho\sigma} + J_{\tau\rho\sigma} \frac{\delta^2(y)}{e_2}, \qquad (3.13)$$

$$\hat{F}_{mn} = F_{mn} - (rg|Q|^2 + \xi)\epsilon_{mn}\frac{\delta^2(y)}{e_2}$$
(3.14)

where  $\xi = \frac{T}{4q}$  is the localized FI term,  $\epsilon_{mn}$  is the 2D volume form and

$$J_{\mu} = \frac{1}{2}i \bigg[ Q^{\dagger} D_{\mu} Q - (D_{\mu} Q)^{\dagger} Q + \frac{1}{2} \bar{\psi}_{Q} \gamma_{\mu} \psi_{Q} - \frac{1}{2} e^{-\frac{1}{2}\phi} \bar{\Lambda} \gamma_{\mu} \Lambda \bigg], \qquad (3.15)$$

$$J_{\tau\rho\sigma} = -\frac{1}{4}\bar{\psi}_Q\gamma_{\tau\rho\sigma}\psi_Q - \frac{1}{8}e^{-\frac{1}{2}\phi}\bar{\Lambda}\gamma_{\tau\rho\sigma}\Lambda.$$
(3.16)

Here in order to cancel the variation of the brane tension action [14], we needed to modify the gauge field strength with the localized FI term<sup>2</sup> proportional to the brane tension. Moreover, the modified field strength for the KR field contains a gauge non-invariant piece proportional to the localized FI term so the gauge transformation of the KR field needs to be modified to (24)

$$\delta_{\Lambda}B_{mn} = \Lambda \left( -\frac{1}{2}F_{mn} + \xi \epsilon_{mn} \frac{\delta^2(y)}{e_2} \right).$$
(3.17)

On the other hand, the SUSY transformations of the bulk fields are the same as eqs. (2.8)–(2.14) with  $G_{MNP}$  and  $F_{MN}$  being replaced by  $\hat{G}_{MNP}$  and  $\hat{F}_{MN}$ , respectively, and the gauge field  $A_M$  being kept the same as in the no-brane case, with an exception that the SUSY transformation of the extra components of the KR field has an additional term as

$$\delta B_{mn} = \frac{1}{4} i \bar{\psi}_Q \varepsilon_+ Q \epsilon_{mn} \frac{\delta^2(y)}{e_2} + \text{h.c..}$$
(3.18)

Furthermore, for the modified field strength tensors, we obtain the Bianchi identities as follows,

$$\partial_{[\mu}\hat{G}_{\nu mn]} = \frac{3}{4}\hat{F}_{[\mu\nu}\hat{F}_{mn]} + \left[\frac{i}{2}(D_{[\mu}Q)^{\dagger}(D_{\nu]}Q) + \frac{1}{4}e|Q|^{2}W_{\mu\nu}\right]\epsilon_{mn}\frac{\delta^{2}(y)}{e_{2}},\qquad(3.19)$$

$$\partial_{[\mu}\hat{F}_{mn]} = -\frac{1}{3}rg\partial_{\mu}|Q|^{2}\epsilon_{mn}\frac{\delta^{2}(y)}{e_{2}}.$$
(3.20)

Then, by using eq. (2.17) with the modified Bianchi identities (3.19) and (3.20), we are able to cancel all the remaining variations of the brane action given in eq. (3.5).

We can extend the result to the more general case with multiple branes. When all the branes preserve the same 4D  $\mathcal{N} = 1$  SUSY, we only have to replace the delta terms appearing in the action and the SUSY/gauge transformations:  $T\delta^2(y)$  with  $\sum_i T_i \delta^2(y-y_i)$ , and  $f(Q)\delta^2(y)$  with  $\sum_i f(Q_i)\delta^2(y-y_i)$ .

<sup>&</sup>lt;sup>2</sup>See ref. [20] for discussion on the FI term in 5D gauged supergravity.

## 3.3 The brane-localized gravitino mass

We introduce a gravitino mass term on the brane. Then, the brane action is supplemented by the supersymmetric gravitino mass terms as

$$\mathcal{L}_{\text{gmass}} = -e_4 \frac{1}{2} W_0 e^{\frac{1}{2}\psi} (\bar{\psi}_{\mu+} \gamma^{\mu\nu} C \bar{\psi}_{\nu+}^T + \bar{\psi}_1 \gamma^{\mu} C \bar{\psi}_{\mu+}^T + \bar{\psi}_2 \gamma^{\mu} C \bar{\psi}_{\mu+}^T + \bar{\lambda}_+ C \bar{\lambda}_+^T) + \text{h.c.} \quad (3.21)$$

where  $W_0$  is a constant parameter and

$$\psi_1 = \psi_{5+} + i\psi_{6+}, \quad \psi_2 = \psi_{5+} - i\psi_{6+}.$$
 (3.22)

We also need to modify the SUSY transformations of the extra components of the gravitino as follows,

$$\delta\psi_{+} = W_{0}e^{\frac{1}{2}\psi}C\bar{\varepsilon}_{+}^{T}\frac{\delta^{2}(y)}{e_{2}},$$
(3.23)

$$\delta\psi_{-} = -W_0 e^{\frac{1}{2}\psi} C \bar{\varepsilon}_{+}^T \frac{\delta^2(y)}{e_2}.$$
 (3.24)

Here  $e^{\psi}$  is the volume modulus of the extra dimensions. Thus, similarly to the ungauged supergravity [15], the brane gravitino mass has a nontrivial coupling to the volume modulus of the extra dimensions. Under the modified gravitino variations (3.23) and (3.24), the variation of the bulk gravitino linear terms would have induced singular terms for a nonzero background gauge flux. So, in order to cancel them, we needed to introduce the branelocalized gaugino mass, which is the same as the gravitino mass. When the superpotential depends on the brane chiral multiplets, we can infer the form of the brane F-term as

$$\mathcal{L}_F = -e_4 e^{\psi - \frac{1}{2}\phi} |F_Q|^2 \tag{3.25}$$

with  $F_Q = \frac{\partial W}{\partial Q}$ . Consequently, similarly to the ungauged supergravity case [15], we show that the F-term has a nontrivial coupling to the dilaton as well as the volume modulus.

#### 4. The supersymmetric brane solution

In the presence of the modification in the gauge field strength in eq. (3.14), the general warped solution obtained in ref. [11, 10, 12] is maintained up to the modified solution for the gauge field [14]. On the other hand, when the KR field and the 4D component of the gauge field are set to zero, the modified field strength for the KR field in eqs. (3.12) and (3.13) does not affect the equations of motion.

Assuming the axial symmetry in the internal space, it has been found that the general warped solution with 4D Minkowski space takes the following form [11, 14],

$$ds^{2} = W^{2}(r)\eta_{\mu\nu}dx^{\mu}dx^{\nu} + R^{2}(r)\bigg(dr^{2} + \lambda^{2}\Theta^{2}(r)d\theta^{2}\bigg), \qquad (4.1)$$

$$F_{mn} = q\epsilon_{mn},\tag{4.2}$$

$$\phi = 4\ln W,\tag{4.3}$$

with

$$R = \frac{W}{f_0}, \quad \Theta = \frac{r}{W^4}, \tag{4.4}$$

$$W^4 = \frac{f_1}{f_0}, \quad f_0 = 1 + \frac{r^2}{r_0^2}, \quad f_1 = 1 + \frac{r^2}{r_1^2},$$
 (4.5)

where q is a constant denoting the magnetic flux, and the two radii  $r_0$ ,  $r_1$  are given by

$$r_0^2 = \frac{1}{2g^2}, \quad r_1^2 = \frac{8}{q^2}.$$
 (4.6)

In the warped solution, the metric has two conical singularities, one at r = 0 and the other at  $r = \infty$ , which is at finite proper distance from the former one. The singular terms coming from the deficit angles  $\delta_i$  at these singularities need to be compensated by brane tensions  $T_i = 2\delta_i (i = 1, 2)$  with the following matching conditions,

$$\frac{\delta_1}{2\pi} = 1 - \lambda,\tag{4.7}$$

$$\frac{\delta_2}{2\pi} = 1 - \lambda \frac{r_1^2}{r_0^2}.$$
(4.8)

Moreover, when the brane actions are invariant under the same 4D  $\mathcal{N} = 1$ , the modified gauge field strength has two singular terms proportional to the brane tensions at the conical singularities,

$$\hat{F}_{mn} = F_{mn} - \epsilon_{mn} \xi_i \frac{\delta^2 (y - y_i)}{e_2} \tag{4.9}$$

with  $\xi_i = \frac{T_i}{4g}$  for i = 1, 2.

We need two patches of coordinates to cover the whole bulk space. In the patch including r = 0, the solution of the only non-zero component of the gauge field is

$$A_{\theta} = -\frac{4\lambda}{q} \left(\frac{1}{f_1} - 1\right) + \frac{\xi_1}{2\pi}.$$
(4.10)

Likewise, the gauge potential in the patch surrounding  $r = \infty$  is

$$A_{\theta} = -\frac{4\lambda}{q} \frac{1}{f_1} - \frac{\xi_2}{2\pi}.$$
(4.11)

Hence, after connecting the gauge field solutions in two patches by a gauge transformation and requiring that it is single valued under  $2\pi$  rotations, we find the following quantization condition should hold

$$\frac{4\lambda g}{q} = n - \frac{g}{2\pi} (\xi_1 + \xi_2), \quad n \in \mathbf{Z}.$$
(4.12)

In other words, we find that the FI terms fix the Wilson line phases of the gauge potential to be non-vanishing on the branes and can contribute to the quantization condition for  $\xi_1 + \xi_2 \neq 0$ , i.e., when  $T_1 + T_2 \neq 0$ . This result can be regarded as a generalization of the 6D global SUSY case discussed in ref. [27] to the case with nonzero bulk gauge flux.

Using the flux quantization (4.12) with eqs. (4.7) and (4.8), we obtain the brane tensions are related as

$$\left(1 - \frac{T_1}{4\pi}\right) \left(1 - \frac{T_2}{4\pi}\right) = \left[n - \frac{g}{2\pi}(\xi_1 + \xi_2)\right]^2.$$
(4.13)

In particular, in the unwarped limit, i.e., for  $r_0 = r_1$ , the solution becomes the football solution with two equal brane tensions  $T_1 = T_2 = 4\pi(1 - \lambda)$ . Then, since q = 4g and  $\xi_1 = \xi_2 = \frac{\pi}{g}(1 - \lambda)$ , the quantization condition (4.12) is satisfied for n = 1 and arbitrary  $\lambda$ . Thus, for  $0 < \lambda < 1$ , brane tensions can be arbitrary and positive. This is a remarkable result, as compared to the non-SUSY brane tension action [9] where the brane tensions are always negative because  $\lambda$  is quantized as a natural number.

Moreover, unlike the general warped solution which breaks SUSY completely, it has been shown [14] that for any  $\lambda$ , the football solution preserves 4D  $\mathcal{N} = 1$  SUSY. For completeness, we add a brief discussion on this result given in ref. [14] in order. In the patch surrounding the brane at r = 0, the nontrivial fermionic SUSY transformations are

$$\delta\lambda = i\sqrt{2}g(\gamma^5 - 1)\varepsilon, \qquad (4.14)$$
$$\delta\psi_\theta = \left[\partial_\theta + \frac{i}{2} \left\{ 1 + \lambda \left(1 - \frac{2}{f_0}\right) \right\} \gamma^5 + i\lambda \left(\frac{1}{f_0} - 1\right) - i\frac{g\xi_0}{2\pi} \right]\varepsilon$$

$$= \left[\partial_{\theta} + \frac{i}{2} \left\{ 1 + \lambda \left( 1 - \frac{2}{f_0} \right) \right\} (\gamma^5 - 1) \right] \varepsilon, \qquad (4.15)$$

where use is made of  $g\xi_0 = \frac{1}{4}T_0 = \pi(1-\lambda)$  from eq. (4.7) in the last line. Then, for a non-zero left-handed variation parameter  $\varepsilon_+$ , for which the gaugino variation is manifestly zero, the remaining nonzero gravitino variation is  $\delta\psi_{\theta+} = \partial_{\theta}\varepsilon_+$ . So, for any  $\lambda$ , i.e. any brane tension, there exists a constant 4D left-handed Killing spinor  $\varepsilon_+$ . Therefore, we can see that the modified gauge potential is crucial for maintaining the 4D  $\mathcal{N} = 1$  SUSY even in the presence of arbitrary brane tensions. It has been also shown that there appears a single chiral massless mode of gravitino for positive brane tensions on the football [14], while there are multiple massless modes of gravitino possible in the case with the non-SUSY brane tension action [16].

## 5. The 4D effective action

We consider the low energy effective action containing light bulk and brane fields after the extra dimensions are compactified on a football. We show that the resulting 4D effective action can be described by a 4D supergravity with the specific form of the Kähler potential containing a constant  $U(1)_R$  FI term, the gauge kinetic functions and the effective superpotential from the branes. We discuss on the effect of the bulk Green-Schwarz term on the 4D anomalies.

#### 5.1 The tree-level effective action

To make a KK dimensional reduction to 4D for the supersymmetric brane solution, we

take the separable ansatz for the  $6D \text{ metric}^3$  as

$$ds^{2} = e^{-\psi(x)}g_{\mu\nu}(x)dx^{\mu}dx^{\nu} + e^{\psi(x)}ds_{2}^{2}$$
(5.1)

where  $ds_2^2$  is the 2D metric of the football solution and  $\psi$  is the volume modulus.

By integrating the supersymmetric bulk-brane system (3.4) over the extra dimensions with eqs. (C.19) and (C.25), the 4D effective action for the bosonic fields apart from the brane F- and D-terms can be obtained as

$$\mathcal{L}_{\text{boson}} = \frac{1}{2} M_*^4 e_4 \int d^2 y e_2 \left[ R(g) - (\partial_\mu \psi)^2 - \frac{1}{4} (\partial_\mu f)^2 - \frac{1}{2} e^{\psi + \frac{1}{2}f} M_*^{-4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{2} e^{2\psi + f} (G_{\mu\nu\rho})^2 - \frac{1}{2} e^{-2\psi + f} \left( \partial_\mu b - \frac{2q}{M_*^4} A_\mu - \frac{i}{M_P^2} (Q^{\dagger} D_\mu Q - (D_\mu Q)^{\dagger} Q) \right)^2 - M_*^{-4} e^{-3\psi + \frac{1}{2}f} \left( q - \frac{rg}{V} |Q|^2 \right)^2 - 4g^2 M_*^4 (-2e^{-2\psi} + e^{-\psi - \frac{1}{2}f}) \right] + e_4 \left[ -e^{-\psi + \frac{1}{2}f} (D^\mu Q)^{\dagger} (D_\mu Q) - \frac{1}{4} W_{\mu\nu} W^{\mu\nu} - 2rg^2 M_*^4 e^{-2\psi} |Q|^2 \right]$$
(5.2)

where  $q = 2gM_*^4$ . Here we have recovered the 6D fundamental scale  $M_*$  that was taken to be  $M_*^4 = 2$ .

We redefine the scalar fields as the mixture of the dilaton and the volume modulus as

$$s = e^{\psi + \frac{1}{2}f}, \quad t = e^{\psi - \frac{1}{2}f}.$$
 (5.3)

We also dualize the 4D component field strength for the KR field as  $e^f G_{\mu\nu\rho} = \epsilon_{\mu\nu\rho\tau} \partial^{\tau} \sigma$ . The effective action (5.2) then becomes

$$\mathcal{L}_{\text{boson}} = M_P^2 \sqrt{-g} \bigg[ \frac{1}{2} R(g) - \frac{(\partial_\mu s)^2}{4s^2} - \frac{(\partial_\mu t)^2}{4t^2} \\ - \frac{1}{4M_P^2} sF_{\mu\nu} F^{\mu\nu} - \frac{1}{M_P^2 t} (D^\mu Q)^\dagger (D_\mu Q) - \frac{1}{4M_P^2} W_{\mu\nu} W^{\mu\nu} \\ - \frac{(\partial_\mu \sigma)^2}{4s^2} - \frac{1}{4t^2} \bigg( \partial_\mu b - 4g_R A_\mu - \frac{i}{M_P^2} (Q^\dagger D_\mu Q - (D_\mu Q)^\dagger Q) \bigg)^2 \\ - \frac{2g_R^2 M_P^2}{s} \bigg\{ 1 - \frac{1}{t} \bigg( 1 - \frac{r}{2M_P^2} |Q|^2 \bigg) \bigg\}^2 \bigg].$$
(5.4)

Here the 4D Planck scale is given by  $M_P^2 = M_*^4 V$  with the volume of the extra dimensions  $V = \lambda \pi r_0^2$ , and the covariant derivative for the brane scalar is given by  $D_\mu Q = (\partial_\mu + irg_R A_\mu - ieW_\mu)Q$  with the 4D effective U(1)<sub>R</sub> gauge coupling  $g_R = g/\sqrt{V}$ . Here we can see that the U(1)<sub>R</sub> gauge boson gets a nonzero mass,  $M_A^2 = 8g_R^2 M_P^2$ , by a Green-Schwarz mechanism due to the nonzero gauge flux.

We now find that the bosonic effective action (5.4) corresponds to the 4D supergravity action with the Kähler potential<sup>4</sup> K containing a constant  $U(1)_R$  FI term, the gauge kinetic

<sup>&</sup>lt;sup>3</sup>We set the shape modulus to zero, i.e.  $\xi = \psi$  in the general metric ansatz in the appendix B, because the shape modulus is massive [21].

<sup>&</sup>lt;sup>4</sup>For simplicity, we omitted the coupling of the brane vector multiplet to the brane chiral multiplet. It can be implemented by replacing  $e^{-2rg_R V_R}$  in eq. (5.5) with  $e^{-2rg_R V_R+2eV_W}$ .

functions for the bulk and brane vector multiplets,  $f_R$  and  $f_W$ :

$$K = -\ln\left(\frac{1}{2}(S+S^{\dagger})\right) - \ln\left(\frac{1}{2}(T+T^{\dagger}-\delta_{GS}V_{R}) - \frac{1}{M_{P}^{2}}\tilde{Q}^{\dagger}e^{-2rg_{R}V_{R}}\tilde{Q}\right) - \frac{2\xi_{R}}{M_{P}^{2}}V_{R}, (5.5)$$
  
$$f_{R} = S, \qquad f_{W} = 1$$
(5.6)

where the scalar components of the moduli supermultiplets are given by

$$S = s + i\sigma, \qquad T = t + \frac{1}{M_P^2} |Q|^2 + ib.$$
 (5.7)

Here the Green-Schwarz parameter is  $\delta_{GS} = 8g_R$  and the constant FI term coefficient is related to the Green-Schwarz parameter by  $\xi_R = \frac{1}{4} \delta_{GS} M_P^2$ . We note that the chiral superfield  $\tilde{Q}$  contains the component fields,  $(Q^*, \psi_Q^c)$ . The nonzero scalar potential in the effective action (5.4) corresponds to the U(1)<sub>R</sub> D-term in the 4D effective supergravity. We can also see that the U(1)<sub>R</sub> gauge kinetic function  $f_R$  has a nontrivial S modulus dependence while the gauge kinetic function  $f_W$  for the brane vector multiplet is trivial.

After making the gravitino kinetic term canonical in the KK reduction to 4D, the effective 4D gravitino mass becomes

$$\mathcal{L}_{\text{gmass}} = -e_4 \frac{1}{2} W_0 e^{-\psi} \bar{\psi}_{\mu+} \gamma^{\mu\nu} C \bar{\psi}_{\nu+}^T + \text{h.c.}.$$
(5.8)

Compared to the gravitino mass in 4D supergravity,  $\mathcal{L}_m = -e_4 \frac{1}{2} e^{K/2} W \bar{\psi}_{\mu+} \gamma^{\mu\nu} C \bar{\psi}_{\nu+}^T + \text{h.c.},$  the effective superpotential is independent of the moduli:

$$W = W_0. \tag{5.9}$$

We note that  $U(1)_R$  gauge invariance requires the superpotential to take an R charge +2 as the gravitino  $\psi_{\mu+}$  has an R charge +1.

We can easily generalize the results to the case when the brane matter is present at the other brane. The 4D effective supergravity is then described by

$$K = -\ln\left(\frac{1}{2}(S+S^{\dagger})\right) - \ln\left(\frac{1}{2}(T+T^{\dagger}-\delta_{GS}V_{R}) - \frac{1}{M_{P}^{2}}\sum_{i=1,2}\tilde{Q}_{i}^{\dagger}e^{-2r_{i}g_{R}V_{R}}\tilde{Q}_{i}\right) - \frac{2\xi_{R}}{M_{P}^{2}}V_{R}, \quad (5.10)$$

$$W = \sum_{i=1,2} W_i(\tilde{Q}_i) \tag{5.11}$$

where the scalar component of the T modulus is generalized to

$$T = t + \frac{1}{M_P^2} \sum_{i=1,2} |Q_i|^2 + ib.$$
(5.12)

Therefore, the scalar potential takes a more general form as

$$V_0 = \frac{2g_R^2 M_P^4}{s} \left[ 1 - \frac{1}{t} \left\{ 1 - \frac{1}{2M_P^2} \left( \sum_{i=1,2} r_i |Q_i|^2 \right) \right\} \right]^2.$$
(5.13)

For a more general Kähler potential for the brane chiral multiplet  $\Omega_i(\tilde{Q}_i^{\dagger}e^{-2rg_R V_R}\tilde{Q}_i)$ , we have to replace the scalar component of the T modulus and the Kähler potential, respectively, by

$$T = t + \sum_{i=1,2} \Omega_i(|Q_i|^2) + ib$$
(5.14)

and

$$K = -\ln\left(\frac{1}{2}(S+S^{\dagger})\right) - \ln\left(\frac{1}{2}(T+T^{\dagger}-\delta_{GS}V_{R}) - \sum_{i=1,2}\Omega_{i}(\tilde{Q}_{i}^{\dagger}e^{-2r_{i}g_{R}V_{R}}\tilde{Q}_{i})\right) - \frac{2\xi_{R}}{M_{P}^{2}}V_{R}.$$
 (5.15)

## **5.2** The $U(1)_R$ gauge transformations

The brane chiral multiplet  $\hat{Q}$  having an R charge r transforms under the  $U(1)_R$  with parameter  $\Phi$  (where  $\operatorname{Re} \Phi|_{\theta=\bar{\theta}=0} = \Lambda_R$ ) as

$$\tilde{Q} \to e^{irg_R \Phi} \tilde{Q} \tag{5.16}$$

while the  $U(1)_R$  vector multiplet transforms as

$$V_R \to V_R + \frac{i}{2}(\Phi - \Phi^{\dagger}).$$
 (5.17)

Gauge invariance of the T-dependent piece of the Kähler potential (5.5) requires that, under the  $U(1)_R$  gauge transformation, the T modulus transforms nonlinearly as

$$T \to T + \frac{i}{2} \delta_{GS} \Phi.$$
 (5.18)

This results in a shift of the axion field,  $b \to b + \frac{1}{2} \delta_{GS} \Lambda_R = b + 4g_R \Lambda_R$ . This is consistent with the globally well-defined KR field [22], given in eq. (C.2),  $\mathcal{B}_{mn} = -b\epsilon_{mn}$ , that has a gauge transformation,  $\delta \mathcal{B}_{mn} = -\Lambda \langle \hat{F}_{mn} \rangle = -2q M_*^{-4} \Lambda \epsilon_{mn}$  with  $\Lambda = \Lambda_R / \sqrt{V}$ . On the other hand, the *S* modulus does not transform under the U(1)<sub>R</sub>.

The constant  $U(1)_R$  FI term appearing in the Kähler potential (5.5) causes the Kähler potential to transform under the  $U(1)_R$  as

$$K \to K - i \frac{\xi_R}{M_P^2} (\Phi - \Phi^{\dagger}).$$
(5.19)

So, the Kähler potential transforms exactly like an abelian vector superfield. In the global SUSY case, such a variation of the Kähler potential would maintain the invariance of the action after the superspace integral. In the local SUSY case, however, the action would not be invariant unless the Weyl rescaling invariance of the supergravity is supersymmetric. In the Weyl compensator formalism [17], the super-Weyl symmetry is manifest due to the chiral compensator superfield. In this case, the 4D supergravity action is written in the superspace form as

$$S = \int d^4x \left[ d^4\theta \,\mathbf{E} \left( -3C^{\dagger}C \,e^{-K/3} \right) + \left( \int d^2\theta \,\mathcal{E}C^3 W + \text{h.c.} \right) \right]$$
(5.20)

where **E** is the full superspace measure,  $\mathcal{E}$  is the chiral superspace measure and C is the chiral compensator superfield. The above action (5.20) is super-Weyl invariant under the following transformations,

$$\mathbf{E} \to e^{2\tau + 2\bar{\tau}} \mathbf{E}, \qquad \mathcal{E} \to e^{6\tau} \mathcal{E},$$
 (5.21)

$$C \to e^{-2\tau} C, \qquad W \to W$$
 (5.22)

with a complex parameter  $\tau$ . On the other hand, we can show that together with eq. (5.19), the action is also invariant under the U(1)<sub>R</sub> transformations,

$$C \to e^{-i\frac{\xi_R}{3M_P^2}\Phi}C,\tag{5.23}$$

$$W \to e^{i\frac{K}{M_P}\Phi}W.$$
(5.24)

Thus, since  $e^{i\xi_R\Phi/M_P^2} = e^{2ig_R\Phi}$ , the effective superpotential takes an R charge +2, as well known in the 4D supergravity with the gauged U(1)<sub>R</sub> [18, 19]. When we choose the super-Weyl gauge  $C = 1 + \theta^2 F_C$  for manifest SUSY and holomorphicity, combining the U(1)<sub>R</sub> transformation and a super-Weyl transformation with  $\tau = -i\frac{\xi_R}{6M_P^2}\Phi$  maintains the super-Weyl gauge while making the superspace action (5.20) gauge invariant. Then, the accompanying super-Weyl transform of the superspace measure **E** and  $\mathcal{E}$  means that the gravitino transforms under the U(1)<sub>R</sub>.

## 5.3 The bulk Green-Schwarz term and the 4D anomalies

In order to cancel the reducible bulk anomalies, it is necessary to introduce a Green-Schwarz(GS) term [23-25] as follows,

$$\mathcal{L}_{GS} = -kvB \wedge \left( \operatorname{tr}(R \wedge R) - \tilde{v}F \wedge F \right)$$
(5.25)

with the extended gauge transformation of the KR field,  $\delta B = -\frac{1}{2}\Lambda F + \omega_L^1/v$ , where  $\delta \omega_L^1 = d\omega_L$  with  $\omega_L$  being the gravitational Chern-Simons(CS) form<sup>5</sup> satisfying  $d\omega_L = \operatorname{tr} R \wedge R$ . Gauge invariance would require to modify the field strength as  $G = dB + \frac{1}{2}F \wedge A - \omega_L/v$ . Here  $k, v, \tilde{v}$  are calculable for the given bulk fermion content.

From eq. (C.2), we can rewrite the KR field in terms of the globally well-defined one as  $B = \mathcal{B} + \frac{1}{2} \langle A \rangle \wedge \mathcal{A}$ . Then, under the gauge transformations,  $\delta_{\Lambda} \mathcal{B} = -\Lambda \langle \hat{F} \rangle$  and  $\delta_{\Lambda} \mathcal{A} = d\Lambda$ , the bulk Green-Schwarz term transforms

$$\delta_{\Lambda} \mathcal{L}_{GS} = -kv \bigg( -\Lambda \langle \hat{F} \rangle + \frac{1}{2} \langle A \rangle \wedge d\Lambda \bigg) \wedge \Big( \operatorname{tr}(R \wedge R) - \tilde{v}F \wedge F \Big).$$
(5.26)

When we focus on the gauge part, the gauge variation of the Green-Schwarz term becomes

$$\delta_{\Lambda} \mathcal{L}_{GS} = -kv\tilde{v} \left[ \frac{1}{2} \Lambda \langle \hat{F} \rangle \wedge F \wedge F + \left( \Lambda \langle A \rangle \wedge F \wedge F \Big|_{r=\infty} - \Lambda \langle A \rangle \wedge F \wedge F \Big|_{r=0} \right) \right]$$
(5.27)

<sup>&</sup>lt;sup>5</sup>Since the GS term and the necessary addition of the gravitational CS term are not invariant under the SUSY transformations, we would need to add more terms in the bulk action. However, considering the complete SUSY action with them is beyond the scope of the paper.

where use is made of eqs. (4.9), (4.10) and (4.11). The first term is a bulk term giving rise to the 4D anomalies induced by the gauge flux. It is also present in the Salam-Sezgin vacuum [8] without brane tensions where there appear 4D chiral massless modes of the bulk fermions even without orbifold projections. On the other hand, the last two terms correspond to the variation of the effective Chern-Simons action [26]. Since  $\langle A \rangle$  does not vanish at the branes due to the localized FI terms, they generate the U(1)<sub>R</sub> gauge anomalies on the boundaries, r = 0 and  $r = \infty$ . The U(1)<sub>R</sub>-mixed gravitational anomalies are also induced both in the bulk and on the boundaries in a similar fashion. Since the localized anomalies are proportional to the arbitrary localized FI terms, in order to cancel the localized anomalies, we need to introduce R charged fermions on the branes. However, since the localized anomalies are restrictive, i.e. the U(1)<sub>R</sub>-mixed gravitational anomalies are proportional to the gauge anomalies depending on the bulk fermion content via  $\tilde{v}$ , the brane fermion content should be constrained unless there are additional localized Green-Schwarz terms [27]. The U(1)<sub>R</sub> anomaly cancellation with/without the Green-Schwarz term has been discussed in 4D supergravity context [28].

## 6. Modulus stabilization and $U(1)_R$ mediation

Although the 4D scalar potential stabilizes one of the moduli, the T modulus, due to a nonzero  $U(1)_R$  D-term, the remaining S modulus needs to be fixed too. We consider a possibility of having the S modulus stabilized by a S-dependent superpotential due to the bulk gaugino condensates. Then, introducing a brane F-term and/or a brane D-term as the uplifting potential, we can have all the moduli stabilized at the Minkowski vacuum. Consequently, due to the shift of the T modulus from the value determined only by the  $U(1)_R$  D-term, we show that the brane scalar gets a nonzero soft mass which is proportional to the R charge of the brane chiral multiplet. We call this kind of mechanism of generating the soft mass " $U(1)_R$  mediation".

## 6.1 Modulus stabilization

From the effective action (5.4), we can read the 4D effective scalar potential as

$$V_0 = \frac{2g_R^2 M_P^4}{s} \left[ 1 - \frac{1}{t} \left( 1 - \frac{r}{2M_P^2} |Q|^2 \right) \right]^2.$$
(6.1)

So, we find that the minimum of the potential occurs at t = 1 and |Q| = 0. Then, at this minimum, the tree-level brane scalar mass vanishes. This is due to the cancellation between the tree-level brane-localized scalar mass term and the flux-induced mass term. Moreover, we note that the vacuum energy at the minimum vanishes without SUSY breakdown. Since the value of the S modulus is not determined, however, there should be additional potential terms that arise in stabilizing the S modulus by some other mechanism.

In order to stabilize the S modulus, we assume that the bulk non-perturbative dynamics generates a modulus potential with S-dependent superpotential W(S). Then, the additional contribution to the 4D scalar potential is

$$V_{1} = \frac{e^{K}}{M_{P}^{2}} \left[ \left| \frac{\partial W}{\partial S} + \frac{\partial K}{\partial S} W \right|^{2} K_{SS^{\dagger}}^{-1} - 2|W|^{2} \right]$$
$$= \frac{1}{M_{P}^{2} st} \left[ (S + S^{\dagger})^{2} \left| \frac{\partial W}{\partial S} - \frac{1}{S + S^{\dagger}} W \right|^{2} - 2|W|^{2} \right].$$
(6.2)

The extremum conditions for the total scalar potential  $V = V_0 + V_1$ ,  $\partial_S V = 0$ ,  $\partial_T V = 0$ and  $\partial_Q V = 0$ , are solved approximately<sup>6</sup> by  $t \simeq 1$ , Q = 0 and S solving  $F_S \simeq 0$ .

For instance, we can consider double gaugino condensates [29] in an extended bulk theory with anomaly-free non-abelian gauge groups.<sup>7</sup> In this case, assuming that the  $U(1)_R$ symmetry is broken spontaneously by the VEV of the *R* charged scalars, the superpotential takes a racetrack form [31]

$$W(S) = \Lambda_1 e^{-\beta_1 S} - \Lambda_2 e^{-\beta_2 S}.$$
(6.3)

Here we ignored a possible modification to the gauge kinetic function due to the supersymmetric completion of the Green-Schwarz term (5.26) and assumed that  $\Lambda_1$  and  $\Lambda_2$  have the R charge +2 due to the presence of matter fields [33]. This is in contrast to the fact that the double gaugino condensates are not possible being consistent with the global  $U(1)_R$  in the heterotic string [32] where the S modulus is shifted by an imaginary amount under the  $U(1)_R$ . The matter fields charged under the condensing gauge groups could give additional contributions to the  $U(1)_R$  D-term but we assumed that an extra singlet having an opposite R charge to the matter fields cancels those contributions.

Then, we find that the  $F_S = 0$  condition is

$$\Lambda_1 e^{-\beta_1 S} (1 + \beta_1 (S + S^{\dagger})) = \Lambda_2 e^{-\beta_2 S} (1 + \beta_2 (S + S^{\dagger})).$$
(6.4)

This fixes both Re S and Im S. When the beta functions of the two condensing gauge groups are similar,  $|\beta_1 - \beta_2| \ll \beta_1$ , the solution to eq. (6.4) occurs at a large Re S where the superpotential description with eq. (6.3) is reliable.

## 6.2 Uplifting and soft masses

After fixing the S modulus, however, the vacuum energy becomes nonzero and negative. Therefore, we need to uplift the vacuum energy to zero. To this purpose, from the 4D reduction of eq. (3.25), we derive a F-term potential at the hidden brane as

$$V_2 = \frac{1}{s} |F_{Q'}|^2. \tag{6.5}$$

Instead of the brane F-term, from the 4D reduction of the brane D-term in eq. (3.4), we can consider a D-term potential at the hidden brane as the uplifting potential,

$$V_3 = \frac{1}{2t^2} D^2. ag{6.6}$$

<sup>&</sup>lt;sup>6</sup>Because of the *T* dependence of the scalar potential coming from the gaugino condensates, the minimum value of *t* is shifted from t = 1. Moreover, due to the *S* dependence of the  $U(1)_R$  D-term, the minimum value of *s* is also shifted compared to the one determined only by  $V_1$ .

<sup>&</sup>lt;sup>7</sup>See ref. [30] for some anomaly-free models containing  $U(1)_R$ .

For generalities, including both brane F- and D-terms for the uplifting potential, the full 4D scalar potential becomes

$$V = V_0 + V_1 + V_2 + V_3$$
  
=  $\frac{2g_R^2 M_P^4}{s} \left[ 1 - \frac{1}{t} \left( 1 - \frac{r}{2M_P^2} |Q|^2 \right) \right]^2$   
+  $\frac{1}{M_P^2 st} \left[ (S + S^{\dagger})^2 \left| \frac{\partial W}{\partial S} - \frac{1}{S + S^{\dagger}} W \right|^2 - 2|W|^2 \right] + \frac{1}{2t^2} D^2 + \frac{1}{s} |F_{Q'}|^2.$  (6.7)

Then,  $\partial_Q V = 0$  is satisfied for Q = 0, which is the minimum for r(t-1) > 0. In the case with Q = 0, the remaining extremum conditions,  $\partial_S V = 0$  and  $\partial_T V = 0$ , are

$$0 = -\frac{4g_R^2 M_P^4}{(S+S^{\dagger})^2} \left(1 - \frac{1}{t}\right)^2 - \frac{2}{(S+S^{\dagger})^2} |F_{Q'}|^2 + \frac{2}{M_P^2 t} \frac{\partial}{\partial S} \left\{ \frac{1}{S+S^{\dagger}} \left[ (S+S^{\dagger})^2 \Big| \frac{\partial W}{\partial s} - \frac{1}{S+S^{\dagger}} W \Big|^2 - 2|W|^2 \right] \right\},$$
(6.8)

$$0 = \frac{4g_R^2 M_P^4}{st^2} \left(1 - \frac{1}{t}\right) - \frac{1}{t^3} D^2 - \frac{1}{M_P^2 st^2} \left[ (S + S^{\dagger})^2 \left| \frac{\partial W}{\partial S} - \frac{1}{S + S^{\dagger}} W \right|^2 - 2|W|^2 \right].$$
(6.9)

Compared to the case without the uplifting potential, from eq. (6.9), the T modulus is shifted to

$$t = \frac{1 + \frac{1}{2}\alpha D^2}{1 - \frac{1}{2}\alpha t V_1} \tag{6.10}$$

where  $\alpha \equiv \frac{s}{2g_R^2 M_P^4}$  and  $V_1$  is the scalar potential coming from the gaugino condensates. Thus, we note that the F-term at the hidden brane does not contribute directly to the shift of the *T* modulus because it is independent of the *T* modulus. On the other hand, the *S* modulus is also determined from eq. (6.8) for the fixed *t*. We take the vacuum energy to be zero by choosing the brane D-term as

$$D^{2} = \frac{-2tV_{1}\left(1 - \frac{1}{4}\alpha tV_{1}\right) - \frac{2}{s}|F_{Q'}|^{2}}{1 + \frac{\alpha}{s}|F_{Q'}|^{2}}.$$
(6.11)

Consequently, after minimizing the moduli potential, we find that a nonzero mass for the brane scalar with nonzero R charge is generated and it is given by the following general formula,

$$m_Q^2 = K_{QQ^{\dagger}}^{-1} \frac{\partial^2 V}{\partial Q \partial Q^{\dagger}} \Big|_{Q=0}$$
  
=  $\frac{2g_R^2 M_P^4}{s} (t-1) \frac{r}{tM_P^2}$   
=  $\frac{D^2 + tV_1}{1 - \frac{1}{2}\alpha tV_1} \frac{\frac{1}{2}r}{tM_P^2}.$  (6.12)

We note that the brane scalar mass does not depend on the  $U(1)_R$  gauge coupling. This is because the effective interaction between the visible and hidden sectors is suppressed by the mass squared of the  $U(1)_R$  vector multiplet,  $M_R^2 = 8g_R^2 M_P^2$ . That is, the gauge coupling dependence is cancelled out by the one of the suppression scale for the  $U(1)_R$  mediation. It would be interesting to compare the  $U(1)_R$  mediation to some relevant results in the literature on the SUSY mediation via an additional anomalous or non-anomalous U(1)gauge field [34-36].

First we consider the case with the brane D-term domination by setting  $F_{Q'} = 0$ . Assuming  $\alpha t |V_1| \ll 1$  and ignoring the vacuum energy contribution due to a nonzero F-term for the S modulus,  $D^2 \simeq -2tV_1$  and  $V_1 \simeq -\frac{2|W|^2}{M_P^2 st}$  so we require the brane D-term to be  $D^2 \simeq \frac{4|W|^2}{M_P^2 s}$ . Then, from eq. (6.12) with  $F_{Q'} = 0$ , the brane scalar mass becomes

$$m_Q^2 \simeq \frac{r}{st} \frac{|W|^2}{M_P^4} = r m_{3/2}^2$$
 (6.13)

where use is made of the gravitino mass given by  $m_{3/2}^2 = e^{K} \frac{|W|^2}{M_P^4} = \frac{1}{st} \frac{|W|^2}{M_P^4}$ . For the positive brane scalar mass squared, we require the *R* charge of the brane scalar to be positive. After integrating out the U(1)<sub>R</sub> vector multiplet, the effective operator for generating the soft mass would be written as  $\int d^4\theta \frac{g_R^2}{M_R^2} W'_{\alpha} W'^{\alpha} \tilde{Q}^{\dagger} \tilde{Q}$  with  $W'_{\alpha}$  being the superfield strength for the hidden sector gauge field. We note, however, that from all the known microscopic models for generating the D-term uplifting potential [35–37], such a large D-term generically gives rise to a very heavy gravitino.

Secondly we take the case with the brane F-term domination for which D = 0. Then, similarly to the D-term domination case, for  $\alpha r |V_1| \ll 1$ , we need the brane F-term to be  $|F_{Q'}|^2 \simeq -4stV_1 \simeq \frac{8|W|^2}{M_P^2}$ . Thus, from eq. (6.12) with D = 0, the brane scalar mass becomes

$$m_Q^2 \simeq -\frac{r}{st} \frac{|W|^2}{M_P^4} = -r m_{3/2}^2.$$
 (6.14)

Thus, the brane scalar mass gets an opposite sign compared to the brane D-term domination, so the R charge of the brane scalar must be negative for the positive scalar mass squared. In the F-term domination, the effective operator for generating the soft mass would be written as  $\int d^4\theta \frac{g_R^2}{M_R^2} Q'^{\dagger} Q' \tilde{Q}^{\dagger} \tilde{Q}$  with Q' being the hidden sector chiral superfield.<sup>8</sup> Therefore, for either brane D- or F-term domination, the tree-level soft mass due to the

Therefore, for either brane D- or F-term domination, the tree-level soft mass due to the  $U(1)_R$  mediation can be positive for the appropriate R charge assignment so that it dominates over the anomaly mediation. Particularly, when the R charges of sleptons are nonzero, we can cure the problem of the negative slepton masses in the anomaly mediation [4].

Before ending the section, we also make a remark on the gaugino masses for the Standard Model gauge group. When the SM gauge fields are localized on the brane, there is no gaugino mass at the tree level because the brane gauge kinetic term is trivial, i.e.  $f_W = 1$ . Thus, one can argue that the gaugino masses are generated at one-loop due to the anomaly mediation [4]. In this case, the gaugino masses are suppressed by the loop factor, compared to the gravitino mass or the scalar mass.

<sup>&</sup>lt;sup>8</sup>Even in the case that the hidden chiral superfield does not have an R charge, it has a gauge coupling to the U(1)<sub>R</sub> vector multiplet as seen from the Kähler potential (5.11).

## 7. Conclusion

We have constructed a consistent SUSY action for brane matter multiplets in a 6D chiral gauged supergravity. Introducing brane chiral multiplets charged under the  $U(1)_R$ , we derived the supersymmetric  $U(1)_R$  coupling to the brane by modifying both the gauge field strength and the field strength for the KR field together with the necessary modifications of the fermionic SUSY transformations. We also notify that the modified field strength for the KR field is consistent with SUSY and  $U(1)_R$  symmetry only at the expense of modifying the SUSY and gauge transformations of the KR field with the singular terms, respectively.

The singular modification of the field strength does not challenge the attempt of finding the meaningful result but rather is necessary for obtaining the consistent 4D effective action. We showed that after a dimensional reduction to 4D on the SUSY football background, the obtained low energy effective action with light bulk and brane modes can be consistently reproduced by the corresponding 4D effective supergravity containing the  $U(1)_R$  gauge symmetry. The resulting  $U(1)_R$  gauge symmetry in the 4D effective theory is anomalous as the bulk Green-Schwarz term generates the 4D anomalies in the bulk and on the boundaries. Moreover, the effective scalar potential coming from the  $U(1)_R$  D-term stabilizes one of two moduli, i.e. the T modulus, at the SUSY Minkowski vacuum. Due to the Green-Schwarz mechanism with the gauge flux, the  $U(1)_R$  gauge symmetry is spontaneously broken by eating up the axionic scalar partner of the T modulus, so the mass of the  $U(1)_R$  gauge field is of order the 4D Planck scale up to the  $U(1)_R$  gauge coupling. A brane scalar with nonzero R charge appears in the  $U(1)_R$  D-term and the mass of the brane scalar vanishes at the minimum of the T modulus due to the cancellation between the *R*-charge dependent brane-localized mass term and the flux-induced mass term. However, the S modulus needs to be stabilized for avoiding an unacceptable effect on the equivalence principle and the cosmology.

For the stabilization of the S modulus, we consider the bulk gaugino condensates in an extended bulk theory with the product of the  $U(1)_R$  and non-abelian gauge groups. Then, since the bulk gauge kinetic function depends linearly on the S modulus, the effective superpotential for the double gaugino condensates takes a racetrack form for the S modulus. In the process of the S modulus stabilization, we introduce brane F- and/or D-term potentials to uplift a negative vacuum energy to zero. Because the scalar potential coming from the gaugino condensates or the brane D-term depends on the T modulus, after the S modulus stabilization, the minimum value of the T modulus is shifted from t = 1that would have been obtained only for the  $U(1)_R$  D-term. Thus, we obtained a nonzero soft mass for the brane scalar proportional to its R charge. For this reason, we owe the obtained SUSY breaking to  $U(1)_R$  mediation. For the R charge of order one, the scalar soft mass is of the same order as the gravitino mass. Depending whether brane F- or D-term dominates in the uplifting potential, the brane scalar soft mass squared can be positive or negative for a fixed R charge. Therefore, according to the nature of the uplifting potential coming from the hidden brane, an appropriate R charge assignment for the brane scalar is required to get a positive scalar soft mass squared.

We have discussed briefly on the  $U(1)_R$  anomalies only from the bulk Green-Schwarz counterterm. Since the integrated 4D  $U(1)_R$  anomalies in our case are constrained in the same way as in 4D supergravity, some important developments in the study of the anomaly-free  $U(1)_R$  symmetry in 4D supergravity [28] would be applicable to our effective 4D supergravity. The 6D gauged supergravity with the consistent  $U(1)_R$  gauge symmetry might have a variety of applications in particle physics and cosmology, for instance, to consider the D-term inflation [19] and explain the baryon/lepton number conservation [38, 28], the  $\mu$  problem [39] and the fermion mass hierarchy in the context of the horizontal U(1) symmetry [40]. We leave the detailed analysis on a realistic model building for one problem or another as a future work.

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## A. Notations and conventions

We use the metric signature (-, +, +, +, +) for the 6D metric. The index conventions are the following: (1) for the Einstein indices we use  $M, N, \dots = 0, \dots, 5, 6$  for the 6D indices,  $\mu, \nu, \dots, = 0, \dots, 3$  for the 4D indices and  $m, n, \dots = 5, 6$  for the internal 2D indices, (2) for the Lorentz indices we use  $A, B, \dots = 0, \dots, 5, 6$  for the 6D indices,  $\alpha, \beta, \dots = 0, \dots, 3$ for the 4D indices and  $a, b, \dots = 5, 6$  for the internal 2D indices.

We take the gamma matrices in the locally flat coordinates, satisfying  $\{\Gamma_A, \Gamma_B\} = 2\eta_{AB}$ , to be

$$\Gamma_{\alpha} = \sigma^1 \otimes \gamma_{\alpha}, \ \ \Gamma_5 = \sigma^1 \otimes \gamma_5, \ \ \Gamma_6 = \sigma^2 \otimes \mathbf{1},$$
 (A.1)

where  $\gamma$ 's are the 4D gamma matrices with  $\gamma_5^2 = 1$  and  $\sigma$ 's are the Pauli matrices with  $[\sigma^i, \sigma^j] = 2i\epsilon_{ijk}\sigma^k$ , with i, j, k = 1, 2, 3,

$$\sigma^{1} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma^{2} = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^{3} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$
 (A.2)

The convention for 4D gamma matrices is that

$$\gamma^{\alpha} = \begin{pmatrix} 0 & \sigma^{\alpha} \\ \bar{\sigma}^{\alpha} & 0 \end{pmatrix}, \quad \gamma^{5} = \begin{pmatrix} \mathbf{1} & 0 \\ 0 & -\mathbf{1} \end{pmatrix}, \tag{A.3}$$

with  $\sigma^{\alpha} = (\mathbf{1}, \sigma^i)$  and  $\bar{\sigma}^{\alpha} = (-\mathbf{1}, \sigma^i)$ . The chirality projection operators are defined as  $P_L = (1 + \gamma^5)/2$  and  $P_R = (1 - \gamma^5)/2$ . We note that  $\Gamma^{\alpha 5} = \mathbf{1} \otimes \gamma^{\alpha} \gamma^5$ ,  $\Gamma^{\alpha 6} = i\sigma^3 \otimes \gamma^{\alpha}$ , and  $\Gamma^{56} = i\sigma^3 \otimes \gamma^5$ . In the text, we use the notation,  $\bar{\chi} = -i\chi^{\dagger}\Gamma^0$ , etc. The curved gamma

matrices on the other hand are given in terms of the ones in the locally flat coordinates as  $\Gamma^M = e_A^M \Gamma^A$  where  $e_A^M$  is the 6D vielbein.

The antisymmetrization of the gamma matrices is defined as

$$\Gamma^{M_1 M_2 \cdots M_n} = \Gamma^{[M_1} \Gamma^{M_2} \cdots \Gamma^{M_n]} = \frac{1}{n!} \sum_p (-1)^p \Gamma^{M_1} \Gamma^{M_2} \cdots \Gamma^{M_n}$$
(A.4)

where  $\sum_p$  is the summation over all permutations. The 6D chirality operator is given by

$$\Gamma_7 = \Gamma_0 \Gamma_1 \cdots \Gamma_6 = \sigma^3 \otimes \mathbf{1}. \tag{A.5}$$

# B. Spin connection

We include the scalar modes in the general warped solution as

$$ds^{2} = e^{-\psi}W^{2}\eta_{\mu\nu}dx^{\mu}dx^{\nu} + e^{\xi}(d\rho^{2} + e^{2(\psi-\xi)}a^{2}d\theta^{2}).$$
(B.1)

Then, the nonzero vielbein components are given by

$$e^{\alpha}_{\mu} = e^{-\frac{1}{2}\psi} W \delta^{\alpha}_{\mu}, \tag{B.2}$$

$$e_m^a = \begin{pmatrix} \cos\theta - \sin\theta\\ \sin\theta \cos\theta \end{pmatrix} \begin{pmatrix} e^{\frac{1}{2}\xi} & 0\\ 0 & ae^{\psi - \frac{1}{2}\xi} \end{pmatrix}.$$
 (B.3)

Therefore, the nonzero components of the spin connection are

$$\omega^{\alpha}{}_{\beta} = \frac{1}{2} (\eta^{\alpha\rho} \eta_{\beta\mu} \partial_{\rho} \psi - \delta^{\rho}_{\beta} \delta^{\alpha}_{\mu} \partial_{\rho} \psi) dx^{\mu}, \tag{B.4}$$

$$\omega^{\alpha}{}_{5} = \cos\theta \left(\frac{W'}{W} - \frac{\psi'}{2}\right) W e^{-\frac{1}{2}(\psi+\xi)} \delta^{\alpha}_{\mu} dx^{\mu} -\frac{1}{2} \cos\theta \eta^{\alpha\beta} \partial_{\beta} \xi W^{-1} e^{\frac{1}{2}(\psi+\xi)} d\rho -\sin\theta \eta^{\alpha\beta} \partial_{\beta} (\psi - \frac{1}{2}\xi) a W^{-1} e^{\frac{1}{2}(3\psi-\xi)} d\theta,$$
(B.5)

$$\omega^{\alpha}{}_{6} = \sin\theta \left(\frac{W'}{W} - \frac{\psi'}{2}\right) W e^{-\frac{1}{2}(\psi+\xi)} \delta^{\alpha}_{\mu} dx^{\mu} -\frac{1}{2} \sin\theta \eta^{\alpha\beta} \partial_{\beta} \xi W^{-1} e^{\frac{1}{2}(\psi+\xi)} d\rho + \cos\theta \eta^{\alpha\beta} \partial_{\beta} (\psi - \frac{1}{2}\xi) a W^{-1} e^{\frac{1}{2}(3\psi-\xi)} d\theta,$$
(B.6)

$$\omega^{5}_{6} = \left[1 - \left(\psi' - \frac{1}{2}\xi' + \frac{a'}{a}\right)ae^{\psi - \xi}\right]d\theta \equiv \omega d\theta \tag{B.7}$$

where prime denotes the derivative with respect to  $\rho$ .

We take the case with  $\xi = \psi$  for which the shape modulus is decoupled [21]. In order to determine the modulus coupling of the brane gravitino mass in section 3.3, the following components of the spin connection contracted with the vielbein,  $\omega_{ABC} \equiv e_A^M \omega_{MBC}$ , were used in the text,

$$\omega_{\alpha 56} = \omega_{5\alpha 6} = 0, \tag{B.8}$$

$$\omega_{5\alpha5} = -\frac{1}{2}\partial_{\alpha}\psi \,e^{\frac{1}{2}\psi}W^{-1} = \omega_{6\alpha6}.\tag{B.9}$$

## C. The effective action with the delta terms

In this section, we present the details for deriving the effective action from the modified field strengths with the delta terms.

## C.1 The globally well-defined KR field

We decompose the gauge field into the background value and the fluctuation as  $A = \langle A \rangle + \mathcal{A}$ . The gauge field has different background values in the different patches, being connected by a background gauge transformation,  $\delta_{\Lambda_0} \langle A \rangle = d\Lambda_0$ , while  $\delta_{\Lambda_0} \mathcal{A} = 0$ . Likewise, the gauge field strength is also decomposed to  $F = dA = \langle F \rangle + d\mathcal{A} \equiv \langle F \rangle + \mathcal{F}$ .

Under the background gauge transformation, the KR field transforms as  $\delta_{\Lambda_0}B_{mn} = \Lambda_0(-\frac{1}{2}F_{mn} + \xi\epsilon_{mn}\frac{\delta^2(y)}{e_2})$ ,  $\delta_{\Lambda_0}B_{m\mu} = -\frac{1}{2}\Lambda_0F_{m\mu}$  and  $\delta_{\Lambda_0}B_{\mu\nu} = -\frac{1}{2}\Lambda_0F_{\mu\nu}$ , so it is not globally well-defined.<sup>9</sup> That is, the background gauge transformation is not of the exact form. The gauge transformation of the derivative of B is given by

$$\delta_{\Lambda_0} dB = -\frac{1}{2} d\Lambda_0 \wedge F$$
  
=  $-\frac{1}{2} d\Lambda_0 \wedge (\langle F \rangle + d\mathcal{A})$   
=  $-\frac{1}{2} d\Lambda_0 \wedge d\mathcal{A} = \frac{1}{2} d(d\Lambda_0 \wedge \mathcal{A}) = \frac{1}{2} \delta_{\Lambda_0} d(\langle A \rangle \wedge \mathcal{A})$  (C.1)

where use is made of  $d\Lambda_0 \wedge \langle F \rangle = 0$ . So, we find that the singular term in the background gauge transform of the KR field does not affect the gauge transform of the derivative of B. Thus, we define the globally well-defined KR field as in the case without the delta term as

$$\mathcal{B} = B - \frac{1}{2} \langle A \rangle \wedge \mathcal{A}. \tag{C.2}$$

Then, we can show that  $\delta_{\Lambda_0} d\mathcal{B} = d(\delta_{\Lambda_0} \mathcal{B}) = 0$ . In this case, the redefined  $\mathcal{B}$  is globally well-defined if there is a solution for a one-form C satisfying

$$\delta \mathcal{B}_{mn} = \Lambda_0 \left( -\frac{1}{2} F_{mn} + \xi \epsilon_{mn} \frac{\delta^2(y)}{e_2} \right) - \frac{1}{2} (d\Lambda_0 \wedge \mathcal{A})_{mn} + (dC)_{mn} = 0, \qquad (C.3)$$

$$\delta \mathcal{B}_{m\mu} = -\frac{1}{2}\Lambda_0 F_{m\mu} - \frac{1}{2}(d\Lambda_0 \wedge \mathcal{A})_{m\mu} + (dC)_{m\mu} = 0, \qquad (C.4)$$

and

$$\delta \mathcal{B}_{\mu\nu} = -\frac{1}{2}\Lambda_0 F_{\mu\nu} - \frac{1}{2}(d\Lambda_0 \wedge \mathcal{A})_{\mu\nu} + (dC)_{\mu\nu} = 0.$$
 (C.5)

Now we consider the general gauge transformation of the derivative of  $\mathcal{B}$  as

$$\delta(\partial_{\mu}\mathcal{B}_{mn} + \partial_{m}\mathcal{B}_{n\mu} + \partial_{n}\mathcal{B}_{\mu m}) = \partial_{\mu}\Lambda\left(-\langle F_{mn}\rangle + \xi\epsilon_{mn}\frac{\delta^{2}(y)}{e_{2}}\right) - \frac{1}{2}\partial_{\mu}\Lambda(\partial_{m}\mathcal{A}_{n} - \partial_{n}\mathcal{A}_{m}).$$
(C.6)

Thus, we obtain

$$\delta \mathcal{B}_{mn} = -\Lambda \langle \hat{F}_{mn} \rangle. \tag{C.7}$$

<sup>&</sup>lt;sup>9</sup>Compared to ref. [22], we also have a singular delta term in the gauge transform.

Moreover, the derivative of the redefined KR field is written in terms of the gauge field background as

$$d\mathcal{B} = dB - \frac{1}{2} \langle F \rangle \wedge \mathcal{A} + \frac{1}{2} \langle A \rangle \wedge d\mathcal{A}.$$
 (C.8)

Therefore, we can rewrite the decomposed field strength for the KR field as

$$G = dB + \frac{1}{2} \langle F \rangle \wedge \mathcal{A} + \frac{1}{2} d\mathcal{A} \wedge \langle A \rangle$$
  
=  $d\mathcal{B} + \langle F \rangle \wedge \mathcal{A}.$  (C.9)

That is, in terms of the components, it is written as

$$G_{\mu m n} = \partial_{\mu} \mathcal{B}_{m n} + \partial_{m} \mathcal{B}_{n \mu} - \partial_{n} \mathcal{B}_{m \mu} + \langle F_{m n} \rangle \mathcal{A}_{\mu}.$$
(C.10)

## C.2 The effective matter coupling of the globally well-defined KR field

We present the details on the effective action coming from the modified field strength for the KR field.

For the globally well-defined KR field  $\mathcal{B}$ , the modified field strength for the KR field is given by

$$\hat{G}_{\mu m n} = \partial_{\mu} \mathcal{B}_{m n} + \partial_{m} \mathcal{B}_{n \mu} - \partial_{n} \mathcal{B}_{m \mu} + \langle F_{m n} \rangle \mathcal{A}_{\mu} + \left( J_{\mu} - \xi \mathcal{A}_{\mu} \right) \epsilon_{m n} \frac{\delta^{2}(y)}{e_{2}}$$
$$= \partial_{\mu} \mathcal{B}_{m n} + \partial_{m} \mathcal{B}_{n \mu} - \partial_{n} \mathcal{B}_{m \mu} + \langle \hat{F}_{m n} \rangle \mathcal{A}_{\mu} + J_{\mu} \epsilon_{m n} \frac{\delta^{2}(y)}{e_{2}}.$$
(C.11)

Restricting ourselves to the football SUSY solution, we assume that the solutions are separable as

$$ds^{2} = e^{-\psi(x)}g_{\mu\nu}(x)dx^{\mu}dx^{\nu} + e^{\psi(x)}ds_{2}^{2}, \qquad (C.12)$$

$$F_{mn} = q\epsilon_{mn}, \quad \phi = f(x) \tag{C.13}$$

where  $ds_2^2$  and  $\epsilon_{mn}$  are the 2D metric and volume form for the static solution. Then, from the equation for the KR field,

$$\partial_M(\sqrt{-g}e^{\phi}\hat{G}^{MNP}) = 0, \qquad (C.14)$$

we obtain the solution,

$$\hat{G}_{\mu m n} = e^{3\psi - f} C_{\mu} \epsilon_{m n} \tag{C.15}$$

with

$$\partial_m C_\mu = 0, \quad \partial_\mu (\sqrt{-g} C^\mu) = 0.$$
 (C.16)

Taking the ansätze,  $\mathcal{B}_{m\mu} = -\epsilon_{mn}\partial^n W_{\mu}$  and  $\mathcal{B}_{mn} = -b(x)\epsilon_{mn}$ , eq. (C.15) with (C.11) becomes

$$\Box^{(2)}W_{\mu} = -\partial_{\mu}b + q\mathcal{A}_{\mu} + J_{\mu}\frac{\delta^{2}(y)}{e_{2}} - e^{3\psi - f}C_{\mu}.$$
 (C.17)

Due to the Stokes theorem for compact dimensions, integration of the left-hand side over the extra dimensions must vanish. Thus, we obtain

$$e^{3\psi-f}C_{\mu} = -\partial_{\mu}b + q\mathcal{A}_{\mu} + \frac{J_{\mu}}{V}$$
(C.18)

where the volume of the extra dimensions  $V = \int d^2 y e_2$ . Therefore, after integrating out the 4D vector component of the globally well-defined KR field, the kinetic term for the KR field with the modified field strength becomes

$$\mathcal{L}_{KR} = -\int d^2 y e_6 \frac{1}{4} e^{\phi} \hat{G}_{\mu m n} \hat{G}^{\mu m n} = -e_4 \frac{V}{2} e^{-2\psi + f} \left( \partial_{\mu} b - q \mathcal{A}_{\mu} - \frac{J_{\mu}}{V} \right)^2.$$
(C.19)

This result is used in the text in section 5.1. We can see that there appears a coupling of the gauge boson to the axion, which is proportional to the flux q. This is nothing but a spontaneously breakdown of the U(1)<sub>R</sub> gauge theory by Green-Schwarz mechanism.

## C.3 The effective action for the $U(1)_R$ D-term for a brane scalar

Next we also consider the details on the effective action coming from the modified gauge field strength.

We first consider the Bianchi identity (3.20) for the modified gauge field strength as

$$\partial_{\mu}\hat{F}_{mn} + \partial_{m}F_{n\mu} - \partial_{n}F_{m\mu} = -rg\partial_{\mu}|Q|^{2}\frac{\delta^{2}(y)}{e_{2}}\epsilon_{mn}.$$
 (C.20)

After decomposing the field strength into the background value and the fluctuation, the modified field strength is given by

$$\hat{F}_{mn} = \langle \hat{F}_{mn} \rangle + \mathcal{F}_{mn} - rg|Q|^2 \frac{\delta^2(y)}{e_2} \epsilon_{mn}.$$
(C.21)

Inserting the above expression into the Bianchi identity, we get the Bianchi identity for the fluctuation as

$$\partial_{\mu}\mathcal{F}_{mn} + \partial_{m}\mathcal{F}_{n\mu} - \partial_{n}\mathcal{F}_{m\mu} = 0. \tag{C.22}$$

In order to cancel the problematic delta term proportional to  $|Q|^2$  in  $\hat{F}_{mn}$ , we take the solution for the fluctuation as

$$\mathcal{F}_{mn} = rg|Q|^2 \left(\frac{\delta^2(y)}{e_2} - \frac{1}{V}\right) \epsilon_{mn}.$$
(C.23)

Here the bulk constant term comes from the requirement that integrating the left-hand side over the extra dimensions for the globally well-defined fluctuation  $\mathcal{A}_m$  vanishes. We can also see that for the given solution for  $A_m$  the Bianchi identity (C.22) can be solved for the globally well-defined  $\mathcal{F}_{m\mu}$ .

Then, inserting the solution (C.23) into eq. (C.21) with  $\langle \hat{F}_{mn} \rangle = q \epsilon_{mn}$ , we obtain the modified field strength as

$$\hat{F}_{mn} = \left(q - \frac{rg|Q|^2}{V}\right)\epsilon_{mn}.$$
(C.24)

Therefore, after integrating over the extra dimensions, the bulk gauge kinetic term for the modified field strength becomes

$$\mathcal{L}_{F} = -\frac{1}{4} \int d^{2}y e_{6} e^{\frac{1}{2}\phi} \hat{F}_{mn} \hat{F}^{mn}$$
  
$$= -e_{4} \frac{V}{2} e^{-3\psi + \frac{1}{2}f} \left(q - \frac{rg|Q|^{2}}{V}\right)^{2}.$$
 (C.25)

This term is part of the  $U(1)_R$  D-term in the 4D effective supergravity that is used in the text in section 5.1.

## References

- [1] H.P. Nilles, Supersymmetry, supergravity and particle physics, Phys. Rept. 110 (1984) 1.
- [2] A.H. Chamseddine, R. Arnowitt and P. Nath, Locally supersymmetric grand unification, Phys. Rev. Lett. 49 (1982) 970;
  H.P. Nilles, Dynamically broken supergravity and the hierarchy problem, Phys. Lett. B 115 (1982) 193;
  H.P. Nilles, M. Srednicki and D. Wyler, Weak interaction breakdown induced by supergravity, Phys. Lett. B 120 (1983) 346;
  R. Barbieri, S. Ferrara and C.A. Savoy, Gauge models with spontaneously broken local supersymmetry, Phys. Lett. B 119 (1982) 343;
  L.J. Hall, J.D. Lykken and S. Weinberg, Supergravity as the messenger of supersymmetry breaking, Phys. Rev. D 27 (1983) 2359;
  N. Ohta, Grand unified theories based on local supersymmetry, Prog. Theor. Phys. 70 (1983) 542.
  [3] M. Dine and A.E. Nelson, Dynamical supersymmetry breaking at low-energies, Phys. Rev. D
- [3] M. Dine and A.E. Nelson, Dynamical supersymmetry breaking at low-energies, Phys. Rev. D 48 (1993) 1277 [hep-ph/9303230];
  M. Dine, A.E. Nelson and Y. Shirman, Low-energy dynamical supersymmetry breaking simplified, Phys. Rev. D 51 (1995) 1362 [hep-ph/9408384];
  M. Dine, A.E. Nelson, Y. Nir and Y. Shirman, New tools for low-energy dynamical supersymmetry breaking, Phys. Rev. D 53 (1996) 2658 [hep-ph/9507378].
- [4] L. Randall and R. Sundrum, Out of this world supersymmetry breaking, Nucl. Phys. B 557 (1999) 79 [hep-th/9810155];
  G.F. Giudice, M.A. Luty, H. Murayama and R. Rattazzi, Gaugino mass without singlets, JHEP 12 (1998) 027 [hep-ph/9810442].
- [5] D.E. Kaplan, G.D. Kribs and M. Schmaltz, Supersymmetry breaking through transparent extra dimensions, Phys. Rev. D 62 (2000) 035010 [hep-ph/9911293];
  Z. Chacko, M.A. Luty, A.E. Nelson and E. Ponton, Gaugino mediated supersymmetry breaking, JHEP 01 (2000) 003 [hep-ph/9911323].
- [6] S. Kachru, R. Kallosh, A. Linde and S.P. Trivedi, De Sitter vacua in string theory, Phys. Rev. D 68 (2003) 046005 [hep-th/0301240];
  C.P. Burgess, R. Kallosh and F. Quevedo, De Sitter string vacua from supersymmetric D-terms, JHEP 10 (2003) 056 [hep-th/0309187];
  K. Choi, A. Falkowski, H.P. Nilles, M. Olechowski and S. Pokorski, Stability of flux compactifications and the pattern of supersymmetry breaking, JHEP 11 (2004) 076 [hep-th/0411066];
  K. Choi, A. Falkowski, H.P. Nilles and M. Olechowski, Soft supersymmetry breaking in KKLT flux compactification, Nucl. Phys. B 718 (2005) 113 [hep-th/0503216];
  - K. Choi, K.S. Jeong and K.-i. Okumura, *Phenomenology of mixed modulus-anomaly mediation in fluxed string compactifications and brane models*, *JHEP* **09** (2005) 039 [hep-ph/0504037];
  - M. Endo, M. Yamaguchi and K. Yoshioka, A bottom-up approach to moduli dynamics in heavy gravitino scenario: superpotential, soft terms and sparticle mass spectrum, Phys. Rev. **D** 72 (2005) 015004 [hep-ph/0504036];

O. Lebedev, H.P. Nilles and M. Ratz, *De Sitter vacua from matter superpotentials*, *Phys. Lett.* **B 636** (2006) 126 [hep-th/0603047].

- [7] K. Choi, SUSY Breaking at the tip of throat and mirage mediation, PoS(STRINGSLHC)022 [arXiv:0705.3330].
- [8] H. Nishino and E. Sezgin, Matter and gauge couplings of N = 2 supergravity in six-dimensions, Phys. Lett. B 144 (1984) 187;
  A. Salam and E. Sezgin, Chiral compactification on Minkowski x S<sup>2</sup> of N = 2 Einstein-Maxwell supergravity in six-dimensions, Phys. Lett. B 147 (1984) 47.
- Y. Aghababaie, C.P. Burgess, S.L. Parameswaran and F. Quevedo, Towards a naturally small cosmological constant from branes in 6D supergravity, Nucl. Phys. B 680 (2004) 389
   [hep-th/0304256].
- [10] Y. Aghababaie et al., Warped brane worlds in six dimensional supergravity, JHEP 09 (2003) 037 [hep-th/0308064];
  C.P. Burgess, F. Quevedo, G. Tasinato and I. Zavala, General axisymmetric solutions and self-tuning in 6D chiral gauged supergravity, JHEP 11 (2004) 069 [hep-th/0408109];
  S.L. Parameswaran, G. Tasinato and I. Zavala, The 6D SuperSwirl, Nucl. Phys. B 737 (2006) 49 [hep-th/0509061].
- [11] G.W. Gibbons, R. Güven and C.N. Pope, 3-branes and uniqueness of the Salam-Sezgin vacuum, Phys. Lett. B 595 (2004) 498 [hep-th/0307238].
- [12] H.M. Lee and C. Lüdeling, The general warped solution with conical branes in six-dimensional supergravity, JHEP 01 (2006) 062 [hep-th/0510026].
- [13] J.-W. Chen, M.A. Luty and E. Ponton, A critical cosmological constant from millimeter extra dimensions, JHEP 09 (2000) 012 [hep-th/0003067];
  S.M. Carroll and M.M. Guica, Sidestepping the cosmological constant with football-shaped extra dimensions, hep-th/0302067;
  I. Navarro, Codimension two compactifications and the cosmological constant problem, JCAP 09 (2003) 004 [hep-th/0302129]; Spheres, deficit angles and the cosmological constant, Class. and Quant. Grav. 20 (2003) 3603 [hep-th/0305014];
  H.-P. Nilles, A. Papazoglou and G. Tasinato, Selftuning and its footprints, Nucl. Phys. B 677 (2004) 405 [hep-th/0309042];
  H.M. Lee, A comment on the self-tuning of cosmological constant with deficit angle on a sphere, Phys. Lett. B 587 (2004) 117 [hep-th/0309050];
  J. Garriga and M. Porrati, Football shaped extra dimensions and the absence of self-tuning, JHEP 08 (2004) 028 [hep-th/0406158].
- [14] H.M. Lee and A. Papazoglou, Supersymmetric codimension-two branes in six-dimensional gauged supergravity, JHEP 01 (2008) 008 [arXiv:0710.4319].
- [15] A. Falkowski, H.M. Lee and C. Lüdeling, Gravity mediated supersymmetry breaking in six dimensions, JHEP 10 (2005) 090 [hep-th/0504091].
- [16] H.M. Lee and A. Papazoglou, Gravitino in six-dimensional warped supergravity, Nucl. Phys. B 792 (2008) 166 [arXiv:0705.1157].
- [17] V. Kaplunovsky and J. Louis, Field dependent gauge couplings in locally supersymmetric effective quantum field theories, Nucl. Phys. B 422 (1994) 57 [hep-th/9402005].

- [18] D.Z. Freedman, Supergravity with axial gauge invariance, Phys. Rev. D 15 (1977) 1173;
  R. Barbieri, S. Ferrara, D.V. Nanopoulos and K.S. Stelle, Supergravity, R invariance and spontaneous supersymmetry breaking, Phys. Lett. B 113 (1982) 219;
  S. Ferrara, L. Girardello, T. Kugo and A. Van Proeyen, Relation between different auxiliary field formulations of N = 1 supergravity coupled to matter, Nucl. Phys. B 223 (1983) 191.
- [19] P. Binetruy, G. Dvali, R. Kallosh and A. Van Proeyen, Fayet-Iliopoulos terms in supergravity and cosmology, Class. and Quant. Grav. 21 (2004) 3137 [hep-th/0402046].
- [20] H. Abe, K. Choi and I.-W. Kim, Fayet-Iliopoulos terms in 5D orbifold supergravity, JHEP 09 (2004) 001 [hep-th/0405100];
  H. Abe and K. Choi, Gauged U(1)<sub>R</sub> symmetries and Fayet-Iliopoulos terms in 5D orbifold supergravity, JHEP 12 (2004) 069 [hep-th/0412174];
  F.P. Correia, M.G. Schmidt and Z. Tavartkiladze, (BPS) Fayet-Iliopoulos terms in 5D orbifold SUGRA, Phys. Lett. B 613 (2005) 83 [hep-th/0410281].
- [21] H.M. Lee and A. Papazoglou, Scalar mode analysis of the warped Salam-Sezgin model, Nucl. Phys. B 747 (2006) 294 [Erratum ibid. B 765 (2007) 200] [hep-th/0602208];
  C.P. Burgess, C. de Rham, D. Hoover, D. Mason and A.J. Tolley, Kicking the rugby ball: perturbations of 6D gauged chiral supergravity, JCAP 02 (2007) 009 [hep-th/0610078].
- [22] A.P. Braun, A. Hebecker and M. Trapletti, Flux stabilization in 6 dimensions: d-terms and loop corrections, JHEP 02 (2007) 015 [hep-th/0611102].
- [23] M.B. Green and J.H. Schwarz, Anomaly cancellation in supersymmetric D = 10 gauge theory and superstring theory, Phys. Lett. **B** 149 (1984) 117.
- [24] J. Erler, Anomaly cancellation in six-dimensions, J. Math. Phys. 35 (1994) 1819 [hep-th/9304104].
- [25] Y. Aghababaie, C.P. Burgess, S.L. Parameswaran and F. Quevedo, SUSY breaking and moduli stabilization from fluxes in gauged 6D supergravity, JHEP 03 (2003) 032 [hep-th/0212091].
- [26] C.G. Callan Jr. and J.A. Harvey, Anomalies and fermion zero modes on strings and domain walls, Nucl. Phys. B 250 (1985) 427.
- [27] H.M. Lee, H.P. Nilles and M. Zucker, Spontaneous localization of bulk fields: the six-dimensional case, Nucl. Phys. B 680 (2004) 177 [hep-th/0309195].
- [28] A.H. Chamseddine and H.K. Dreiner, Anomaly free gauged R symmetry in local supersymmetry, Nucl. Phys. B 458 (1996) 65 [hep-ph/9504337];
  D.J. Castano, D.Z. Freedman and C. Manuel, Consequences of supergravity with gauged U(1)<sub>R</sub> symmetry, Nucl. Phys. B 461 (1996) 50 [hep-ph/9507397].
- [29] J.P. Derendinger, L.E. Ibáñez and H.P. Nilles, On the low-energy D = 4, N = 1 supergravity theory extracted from the D = 10, N = 1 superstring, Phys. Lett. B 155 (1985) 65;
  M. Dine, R. Rohm, N. Seiberg and E. Witten, Gluino condensation in superstring models, Phys. Lett. B 156 (1985) 55.
- [30] S. Randjbar-Daemi, A. Salam, E. Sezgin and J.A. Strathdee, An anomaly free model in six-dimensions, Phys. Lett. B 151 (1985) 351;
  S.D. Avramis, A. Kehagias and S. Randjbar-Daemi, A new anomaly-free gauged supergravity in six dimensions, JHEP 05 (2005) 057 [hep-th/0504033];

S.D. Avramis and A. Kehagias, A systematic search for anomaly-free supergravities in six dimensions, JHEP 10 (2005) 052 [hep-th/0508172];

R. Suzuki and Y. Tachikawa, More anomaly-free models of six-dimensional gauged supergravity, J. Math. Phys. 47 (2006) 062302 [hep-th/0512019].

- [31] N.V. Krasnikov, On supersymmetry breaking in superstring theories, Phys. Lett. B 193 (1987) 37.
- [32] A.E. Nelson and N. Seiberg, R symmetry breaking versus supersymmetry breaking, Nucl. Phys. B 416 (1994) 46 [hep-ph/9309299].
- [33] A. Achucarro, B. de Carlos, J.A. Casas and L. Doplicher, De Sitter vacua from uplifting D-terms in effective supergravities from realistic strings, JHEP 06 (2006) 014 [hep-th/0601190].
- [34] Y. Kawamura and T. Kobayashi, Soft scalar masses in string models with anomalous U(1) symmetry, Phys. Lett. B 375 (1996) 141 [Erratum ibid. B 388 (1996) 867] [hep-ph/9601365];

P. Binetruy and E. Dudas, Gaugino condensation and the anomalous U(1), Phys. Lett. B **389** (1996) 503 [hep-th/9607172];

G.R. Dvali and A. Pomarol, Anomalous U(1) as a mediator of supersymmetry breaking, Phys. Rev. Lett. 77 (1996) 3728 [hep-ph/9607383];

Z. Lalak, Anomalous D-term, dynamical supersymmetry breaking and dynamical gauge couplings, Nucl. Phys. B 521 (1998) 37 [hep-ph/9708410];

N. Arkani-Hamed, M. Dine and S.P. Martin, *Dynamical supersymmetry breaking in models with a Green-Schwarz mechanism*, *Phys. Lett.* B 431 (1998) 329 [hep-ph/9803432];

T. Barreiro, B. de Carlos, J.A. Casas and J.M. Moreno, Anomalous U(1), gaugino

condensation and supergravity, Phys. Lett. B 445 (1998) 82 [hep-ph/9808244];

N. Kitazawa, N. Maru and N. Okada, Models of dynamical supersymmetry breaking with gauged U(1)<sub>R</sub> symmetry, Nucl. Phys. B 586 (2000) 261 [hep-ph/0003240]; R-mediation of dynamical supersymmetry breaking, Phys. Rev. D 63 (2001) 015005 [hep-ph/0007253];
F. Brummer, A. Hebecker and M. Trapletti, SUSY breaking mediation by throat fields, Nucl. Phys. B 755 (2006) 186 [hep-th/0605232];

C.A. Scrucca, Soft masses in superstring models with anomalous U(1) symmetries, JHEP 12 (2007) 092 [arXiv:0710.5105];

E. Dudas, Y. Mambrini, S. Pokorski and A. Romagnoni, *Moduli stabilization with Fayet-Iliopoulos uplift*, *JHEP* **04** (2008) 015 [arXiv:0711.4934];

P. Langacker, G. Paz, L.-T. Wang and I. Yavin, Z'-mediated supersymmetry breaking, Phys. Rev. Lett. **100** (2008) 041802 [arXiv:0710.1632]; Aspects of Z'-mediated supersymmetry breaking, arXiv:0801.3693;

R. Dermisek, H. Verlinde and L.-T. Wang, *Hypercharged anomaly mediation*, *Phys. Rev. Lett.* **100** (2008) 131804 [arXiv:0711.3211].

- [35] E. Dudas and S.K. VemPati, Large D-terms, hierarchical soft spectra and moduli stabilisation, Nucl. Phys. B 727 (2005) 139 [hep-th/0506172].
- [36] K. Choi and K.S. Jeong, Supersymmetry breaking and moduli stabilization with anomalous U(1) gauge symmetry, JHEP 08 (2006) 007 [hep-th/0605108].
- [37] H. Jockers and J. Louis, D-terms and F-terms from D7-brane fluxes, Nucl. Phys. B 718 (2005) 203 [hep-th/0502059];
  G. Villadoro and F. Zwirner, De Sitter vacua via consistent D-terms, Phys. Rev. Lett. 95 (2005) 231602 [hep-th/0508167];

E. Dudas and Y. Mambrini, *Moduli stabilization with positive vacuum energy*, *JHEP* **10** (2006) 044 [hep-th/0607077];

M. Haack, D. Krefl, D. Lüst, A. Van Proeyen and M. Zagermann, *Gaugino condensates and D-terms from D7-branes, JHEP* **01** (2007) 078 [hep-th/0609211].

- [38] W. Buchmüller and D. Wyler, *CP violation and R invariance in supersymmetric models of strong and electroweak interactions, Phys. Lett.* B 121 (1983) 321.
- [39] J.E. Kim and H.P. Nilles, The μ problem and the strong CP problem, Phys. Lett. B 138 (1984) 150;
  G.F. Giudice and A. Masiero, A natural solution to the μ problem in supergravity theories, Phys. Lett. B 206 (1988) 480;
  E.J. Chun, Gauged R symmetry, fermion and Higgs mass problem, Phys. Lett. B 367 (1996) 226 [hep-ph/9508320];

H.-P. Nilles and N. Polonsky, Supersymmetric neutrino masses, R-symmetries and the generalized  $\mu$  problem, Nucl. Phys. B 484 (1997) 33 [hep-ph/9606388];

K. Choi, E.J. Chun and H.D. Kim, The  $\mu$  problem and B and L conservation with a discrete gauge R symmetry, Phys. Rev. D 55 (1997) 7010 [hep-ph/9610504].

[40] L.E. Ibáñez and G.G. Ross, Fermion masses and mixing angles from gauge symmetries, Phys. Lett. B 332 (1994) 100 [hep-ph/9403338];
P. Binetruy and P. Ramond, Yukawa textures and anomalies, Phys. Lett. B 350 (1995) 49

[hep-ph/9412385];

V. Jain and R. Shrock, Models of fermion mass matrices based on a flavor dependent and generation dependent U(1) gauge symmetry, Phys. Lett. B 352 (1995) 83 [hep-ph/9412367];
Y. Nir, Gauge unification, Yukawa hierarchy and the μ problem, Phys. Lett. B 354 (1995) 107 [hep-ph/9504312].